# empymod Documentation

Release 1.8.3

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## Version: 1.8.3 ~ Date: 05 July 2019

The electromagnetic modeller **empymod** can model electric or magnetic responses due to a three-dimensional electric or magnetic source in a layered-earth model with vertical transverse isotropic (VTI) resistivity, VTI electric permittivity, and VTI magnetic permeability, from very low frequencies (DC) to very high frequencies (GPR). The calculation is carried out in the wavenumber-frequency domain, and various Hankel- and Fourier-transform methods are included to transform the responses into the space-frequency and space-time domains.

# CHAPTER 1

# More information

For more information regarding installation, usage, contributing, roadmap, bug reports, and much more, see

- Website: https://empymod.github.io,
- Documentation: https://empymod.readthedocs.io,
- Source Code: https://github.com/empymod,
- Examples: https://github.com/empymod/empymod-examples.

# CHAPTER 2

# Features

- Calculates the complete (diffusion and wave phenomena) 3D electromagnetic field in a layered-earth model including vertical transverse isotropic (VTI) resistivity, VTI electric permittivity, and VTI magnetic permeability, for electric and magnetic sources as well as electric and magnetic receivers.
- Modelling routines:
  - bipole: arbitrary oriented, finite length bipoles with given source strength; space-frequency and space-time domains.
  - dipole: infinitesimal small dipoles oriented along the principal axes, normalized field; space-frequency and space-time domains.
  - wavenumber: as dipole, but returns the wavenumber-frequency domain response.
  - gpr: calculates the ground-penetrating radar response for given central frequency, using a Ricker wavelet (experimental).
  - analytical: interface to the analytical, space-frequency and space-time domain solutions.
- Hankel transforms (wavenumber-frequency to space-frequency transform):
  - Digital Linear Filters DLF (using included filters or providing own ones)
  - Quadrature with extrapolation QWE
  - Adaptive quadrature QUAD
- Fourier transforms (space-frequency to space-time transform): Digital Linear Filters DLF (using included filters or providing own ones) Quadrature with extrapolation QWE Logarithmic Fast Fourier Transform FFTLog Fast Fourier Transform FFT
- Analytical, space-frequency and space-time domain solutions:
  - Complete full-space (electric and magnetic sources and receivers); space-frequency domain
  - Diffusive half-space (electric sources and receivers); space-frequency and space-time domains:
    - \* Direct wave (= diffusive full-space solution)
    - \* Reflected wave
    - \* Airwave (semi-analytical in the case of step responses)
- Add-ons (empymod.scripts):

The add-ons for empymod provide some very specific, additional functionalities:

- tmtemod: Return up- and down-going TM/TE-mode contributions for x-directed electric sources and receivers, which are located in the same layer.
- fdesign: Design digital linear filters for the Hankel and Fourier transforms.

# CHAPTER $\mathbf{3}$

Installation

# You can install empymod either via conda:

conda install -c prisae empymod

or via pip:

pip install empymod

Required are Python version 3.5 or higher and the modules *NumPy* and *SciPy*. Consult the installation notes in the manual for more information regarding installation and requirements.

# CHAPTER 4

Citation

If you publish results for which you used empymod, please give credit by citing Werthmüller (2017):

Werthmüller, D., 2017, An open-source full 3D electromagnetic modeler for 1D VTI media in Python: empymod: Geophysics, 82(6), WB9–WB19; DOI: 10.1190/geo2016-0626.1.

All releases have a Zenodo-DOI, provided on the release-page. Also consider citing Hunziker et al. (2015) and Key (2012), without which empymod would not exist.

# CHAPTER 5

# License information

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# 5.1 Manual

# 5.1.1 Theory

The code is principally based on

- [HuTS15] for the wavenumber-domain calculation (kernel),
- [Key12] for the DLF and QWE transforms,
- [SIHM10] for the analytical half-space solutions, and
- [Hami00] for the FFTLog.

See these publications and all the others given in the *References*, if you are interested in the theory on which empymod is based. Another good reference is [ZiS119]. The book derives in great detail the equations for layered-Earth CSEM modelling.

# 5.1.2 Installation

You can install empymod either via conda:

```
conda install -c prisae empymod
```

or via pip:

pip install empymod

Required are Python version 3.5 or higher and the modules NumPy and SciPy. The module numexpr is required additionally (built with Intel's VML) if you want to run parts of the kernel in parallel.

The modeller empymod comes with add-ons (empymod.scripts). These add-ons provide some very specific, additional functionalities. Some of these add-ons have additional, optional dependencies for other modules such as matplotlib. See the *Add-ons*-section for their documentation.

If you are new to Python I recommend using a Python distribution, which will ensure that all dependencies are met, specifically properly compiled versions of NumPy and SciPy; I recommend using Anaconda. If you install Anaconda you can simply start the *Anaconda Navigator*, add the channel prisae and empymod will appear in the package list and can be installed with a click.

The structure of empymod is:

- model.py: EM modelling routines.
- utils.py: Utilities for model such as checking input parameters.
- kernel.py: Kernel of empymod, calculates the wavenumber-domain electromagnetic response. Plus analytical, frequency-domain full- and half-space solutions.
- **transform.py**: Methods to carry out the required Hankel transform from wavenumber to space domain and Fourier transform from frequency to time domain.
- filters.py: Filters for the Digital Linear Filters method DLF (Hankel and Fourier transforms).

# 5.1.3 Usage/Examples

A good starting point is [Wert17b], and more information can be found in [Wert17]. There are a lot of examples of its usage available, in the form of Jupyter notebooks. Have a look at the following repositories:

- Example notebooks: https://github.com/empymod/empymod-examples,
- · Geophysical Tutoriol TLE: https://github.com/empymod/article-tle2017, and
- Numerical examples of [ZiS119]: https://github.com/empymod/csem-ziolkowski-and-slob.

The main modelling routines is bipole, which can calculate the electromagnetic frequency- or time-domain field due to arbitrary finite electric or magnetic bipole sources, measured by arbitrary finite electric or magnetic bipole receivers. The model is defined by horizontal resistivity and anisotropy, horizontal and vertical electric permittivities and horizontal and vertical magnetic permeabilities. By default, the electromagnetic response is normalized to source and receiver of 1 m length, and source strength of 1 A.

A simple frequency-domain example, with most of the parameters left at the default value:

```
>>> import numpy as np
>>> from empymod import bipole
>>> # x-directed bipole source: x0, x1, y0, y1, z0, z1
>>> src = [-50, 50, 0, 0, 100, 100]
>>> # x-directed dipole source-array: x, y, z, azimuth, dip
>>> rec = [np.arange(1, 11) *500, np.zeros(10), 200, 0, 0]
>>> # layer boundaries
>>> depth = [0, 300, 1000, 1050]
>>> # layer resistivities
>>> res = [1e20, .3, 1, 50, 1]
>>> # Frequency
>>> freq = 1
>>> # Calculate electric field due to an electric source at 1 Hz.
>>> # [msrc = mrec = True (default)]
>>> EMfield = bipole(src, rec, depth, res, freq, verb=4)
:: empymod START ::
          [m] : 0 300 1000 1050
  depth
  res
          [Ohm.m] : 1E+20 0.3 1 50 1
          [-]: 11111
  aniso
              [-] :
  epermH
                     1 1 1 1 1
             [-]: 1 1 1 1 1
  epermV
```

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```
[-]: 11111
  mpermH
               [-]: 11111
  mpermV
  frequency [Hz] : 1
  Hankel : DLF (Fast Hankel Transform)
> Filter : Key 201 (2009)
    > Filter : Key 201 (2009)
> DLF type : Standard
  Kernel Opt. : None
   Loop over
                   : None (all vectorized)
    ource(s) : 1 bipole(s)
> intpts : 1 (as dipole)
> length [m] : 100
> x c [m] · 0
   Source(s)

[111]:
y_C [m]:
> z_C [m']
> ac'

     > x_c [m] :
                       0
                       0
                [m] : 100
     > azimuth [°] : 0
     > dip [°] : 0
                  : 10 dipole(s)
   Receiver(s)
     > x [m] : 500 - 5000 : 10 [min-max; #]
                    : 500 1000 1500 2000 2500 3000 3500 4000 4500 5000
     > Y
               [m] : 0 - 0 : 10 [min-max; #]
                    : 0 0 0 0 0 0 0 0 0 0
     > z [m] : 200
     > azimuth [°] : 0
     > dip [°] : 0
   Required ab's : 11
:: empymod END; runtime = 0:00:00.005536 :: 1 kernel call(s)
>>> print (EMfield)
[ 1.68809346e-10 -3.08303130e-10j -8.77189179e-12 -3.76920235e-11j
  -3.46654704e-12 -4.87133683e-12j -3.60159726e-13 -1.12434417e-12j
   1.87807271e-13 -6.21669759e-13j 1.97200208e-13 -4.38210489e-13j
1.44134842e-13 -3.17505260e-13j 9.92770406e-14 -2.33950871e-13j
   6.75287598e-14 -1.74922886e-13j 4.62724887e-14 -1.32266600e-13j]
```

## Hook for user-defined calculation of $\eta$ and $\zeta$

In principal it is always best to write your own modelling routine if you want to adjust something. Just copy empymod.dipole or empymod.bipole as a template, and modify it to your needs. Since empymod v1. 7.4, however, there is a hook which allows you to modify  $\eta_h$ ,  $\eta_v$ ,  $\zeta_h$ , and  $\zeta_v$  quite easily.

The trick is to provide a dictionary (we name it inp here) instead of the resistivity vector in res. This dictionary, inp, has two mandatory plus optional entries:

- res: the resistivity vector you would have provided normally (mandatory).
- A function name, which has to be either or both of (mandatory)
  - func\_eta: To adjust etaH and etaV, or
  - func\_zeta: to adjust zetaH and zetaV.
- In addition, you have to provide all parameters you use in func\_eta/func\_zeta and are not already provided to empymod. All additional parameters must have #layers elements.

The functions func\_eta and func\_zeta must have the following characteristics:

- The signature is func(inp, p\_dict), where
  - inp is the dictionary you provide, and
  - p\_dict is a dictionary that contains all parameters so far calculated in empymod [locals()].
- It must return etaH, etaV if func\_eta, or zetaH, zetaV if func\_zeta.

**Dummy example** 

```
def my_new_eta(inp, p_dict):
    # Your calculations, using the parameters you provided
    # in `inp` and the parameters from empymod in `p_dict`.
    # In the example line below, we provide, e.g., inp['tau']
    return etaH, etaV
```

And then you call empymod with res={'res': res-array, 'tau': tau, 'func\_eta': my\_new\_eta}.

Have a look at the example 2d\_Cole-Cole-IP in the empymod-examples repository, where this hook is exploited in the low-frequency range to use the Cole-Cole model for IP calculation. It could also be used in the high-frequency range to model dielectricity.

## 5.1.4 Contributing

New contributions, bug reports, or any kind of feedback is always welcomed! Have a look at the Roadmap-section to get an idea of things that could be implemented. The best way for interaction is at https://github.com/empymod. If you prefer to contact me outside of GitHub use the contact form on my personal website, https://werthmuller.org.

To install empymod from source, you can download the latest version from GitHub and either add the path to empymod to your python-path variable, or install it in your python distribution via:

python setup.py install

Please make sure your code follows the pep8-guidelines by using, for instance, the python module flake8, and also that your code is covered with appropriate tests. Just get in touch if you have any doubts.

## 5.1.5 Tests and benchmarks

The modeller comes with a test suite using pytest. If you want to run the tests, just install pytest and run it within the empymod-top-directory.

```
> pip install pytest coveralls pytest-flake8 pytest-mpl
> # and then
> cd to/the/empymod/folder # Ensure you are in the right directory,
> ls -d */ # your output should look the same.
docs/ empymod/ tests/
> # pytest will find the tests, which are located in the tests-folder.
> # simply run
> pytest --cov=empymod --flake8 --mpl
```

It should run all tests successfully. Please let me know if not!

Note that installations of empymod via conda or pip do not have the test-suite included. To run the test-suite you must download empymod from GitHub.

There is also a benchmark suite using *airspeed velocity*, located in the empymod/empymod-asv-repository. The results of my machine can be found in the empymod/empymod-bench, its rendered version at empymod.github.io/empymod-asv.

## 5.1.6 Transforms

Included Hankel transforms:

- Digital Linear Filters DLF
- Quadrature with Extrapolation QWE
- Adaptive quadrature QUAD

Included Fourier transforms:

- Digital Linear Filters DLF
- Quadrature with Extrapolation QWE
- Logarithmic Fast Fourier Transform FFTLog
- Fast Fourier Transform FFT

## **Digital Linear Filters**

The module empymod.filters comes with many DLFs for the Hankel and the Fourier transform. If you want to export one of these filters to plain ascii files you can use the tofile-routine of each filter:

```
>>> import empymod
>>> # Load a filter
>>> filt = empymod.filters.wer_201_2018()
>>> # Save it to pure ascii-files
>>> filt.tofile()
>>> # This will save the following three files:
>>> # ./filters/wer_201_2018_base.txt
>>> # ./filters/wer_201_2018_j0.txt
>>> # ./filters/wer_201_2018_j1.txt
```

Similarly, if you want to use an own filter you can do that as well. The filter base and the filter coefficient have to be stored in separate files:

```
>>> import empymod
>>> # Create an empty filter;
>>> # Name has to be the base of the text files
>>> filt = empymod.filters.DigitalFilter('my-filter')
>>> # Load the ascii-files
>>> filt.fromfile()
>>> # This will load the following three files:
>>> # ./filters/my-filter_base.txt
>>> # ./filters/my-filter_j0.txt
>>> # ./filters/my-filter_j1.txt
>>> # and store them in filt.base, filt.j0, and filt.j1.
```

The path can be adjusted by providing tofile and fromfile with a path-argument.

## FFTLog

FFTLog is the logarithmic analogue to the Fast Fourier Transform FFT originally proposed by [Talm78]. The code used by empymod was published in Appendix B of [Hami00] and is publicly available at casa.colorado.edu/~ajsh/FFTLog. From the FFTLog-website:

*FFTLog is a set of fortran subroutines that compute the fast Fourier or Hankel (= Fourier-Bessel) transform of a periodic sequence of logarithmically spaced points.* 

FFTlog can be used for the Hankel as well as for the Fourier Transform, but currently empymod uses it only for the Fourier transform. It uses a simplified version of the python implementation of FFTLog, pyfftlog (github.com/prisae/pyfftlog).

[HaJo88] proposed a logarithmic Fourier transform (abbreviated by the authors as LFT) for electromagnetic geophysics, also based on [Talm78]. I do not know if Hamilton was aware of the work by Haines and Jones. The two publications share as reference only the original paper by Talman, and both cite a publication of Anderson; Hamilton cites [Ande82], and Haines and Jones cite [Ande79]. Hamilton probably never heard of Haines and Jones, as he works in astronomy, and Haines and Jones was published in the *Geophysical Journal*.

Logarithmic FFTs are not widely used in electromagnetics, as far as I know, probably because of the ease, speed, and generally sufficient precision of the digital filter methods with sine and cosine transforms ([Ande75]). However, comparisons show that FFTLog can be faster and more precise than digital filters, specifically for responses with source and receiver at the interface between air and subsurface. Credit to use FFTLog in electromagnetics goes to David Taylor who, in the mid-2000s, implemented FFTLog into the forward modellers of the company Multi-Transient ElectroMagnetic (MTEM Ltd, later Petroleum Geo-Services PGS). The implementation was driven by land responses, where FFTLog can be much more precise than the filter method for very early times.

### **Notes on Fourier Transform**

The Fourier transform to obtain the space-time domain impulse response from the complex-valued spacefrequency response can be calculated by either a cosine transform with the real values, or a sine transform with the imaginary part,

$$\begin{split} E(r,t)^{\text{Impulse}} &= \frac{2}{\pi} \int_0^\infty \Re[E(r,\omega)] \, \cos(\omega t) \, \mathrm{d}\omega \;, \\ &= -\frac{2}{\pi} \int_0^\infty \Im[E(r,\omega)] \, \sin(\omega t) \, \mathrm{d}\omega \;, \end{split}$$

see, e.g., [Ande75] or [Key12]. Quadrature-with-extrapolation, FFTLog, and obviously the sine/cosine-transform all make use of this split.

To obtain the step-on response the frequency-domain result is first divided by  $i\omega$ , in the case of the step-off response it is additionally multiplied by -1. The impulse-response is the time-derivative of the step-response,

$$E(r,t)^{\text{Impulse}} = \frac{\partial E(r,t)^{\text{step}}}{\partial t}$$

Using  $\frac{\partial}{\partial t} \Leftrightarrow i\omega$  and going the other way, from impulse to step, leads to the divison by  $i\omega$ . (This only holds because we define in accordance with the causality principle that  $E(r, t \leq 0) = 0$ ).

With the sine/cosine transform (ft='ffht'/'sin'/'cos') you can choose which one you want for the impulse responses. For the switch-on response, however, the sine-transform is enforced, and equally the cosine transform for the switch-off response. This is because these two do not need to now the field at time 0, E(r, t = 0).

The Quadrature-with-extrapolation and FFTLog are hard-coded to use the cosine transform for step-off responses, and the sine transform for impulse and step-on responses. The FFT uses the full complex-valued response at the moment.

For completeness sake, the step-on response is given by

$$E(r,t)^{\text{Step-on}} = -\frac{2}{\pi} \int_0^\infty \Im\left[\frac{E(r,\omega)}{i\omega}\right] \,\sin(\omega t) \,\mathrm{d}\omega \;,$$

and the step-off by

$$E(r,t)^{\text{Step-off}} = -\frac{2}{\pi} \int_0^\infty \Re\left[\frac{E(r,\omega)}{i\omega}\right] \cos(\omega t) \,\mathrm{d}\omega$$

## 5.1.7 Note on speed, memory, and accuracy

There is the usual trade-off between speed, memory, and accuracy. Very generally speaking we can say that the *DLF* is faster than *QWE*, but *QWE* is much easier on memory usage. *QWE* allows you to control the accuracy. A standard quadrature in the form of *QUAD* is also provided. *QUAD* is generally orders of magnitudes slower, and more fragile depending on the input arguments. However, it can provide accurate results where *DLF* and *QWE* fail.

Parts of the kernel can run in parallel using *numexpr*. This option is activated by setting opt='parallel' (see subsection *Parallelisation*). It is switched off by default.

## Memory

By default empymod will try to carry out the calculation in one go, without looping. If your model has many offsets and many frequencies this can be heavy on memory usage. Even more so if you are calculating time-domain responses for many times. If you are running out of memory, you should use either loop='off' or loop='freq' to loop over offsets or frequencies, respectively. Use verb=3 to see how many offsets and how many frequencies are calculated internally.

## Depths, Rotation, and Bipole

**Depths**: Calculation of many source and receiver positions is fastest if they remain at the same depth, as they can be calculated in one kernel-call. If depths do change, one has to loop over them. Note: Sources or receivers placed on a layer interface are considered in the upper layer.

**Rotation**: Sources and receivers aligned along the principal axes x, y, and z can be calculated in one kernel call. For arbitrary oriented di- or bipoles, 3 kernel calls are required. If source and receiver are arbitrary oriented, 9 (3x3) kernel calls are required.

**Bipole**: Bipoles increase the calculation time by the amount of integration points used. For a source and a receiver bipole with each 5 integration points you need 25 (5x5) kernel calls. You can calculate it in 1 kernel call if you set both integration points to 1, and therefore calculate the bipole as if they were dipoles at their centre.

**Example**: For 1 source and 10 receivers, all at the same depth, 1 kernel call is required. If all receivers are at different depths, 10 kernel calls are required. If you make source and receivers bipoles with 5 integration points, 250 kernel calls are required. If you rotate the source arbitrary horizontally, 500 kernel calls are required. If you rotate the receivers too, in the horizontal plane, 1'000 kernel calls are required. If you rotate the receivers also vertically, 1'500 kernel calls are required. If you rotate the source vertically too, 2'250 kernel calls are required. So your calculation will take 2'250 times longer! No matter how fast the kernel is, this will take a long time. Therefore carefully plan how precise you want to define your source and receiver bipoles.

	source bipole			receiver bipole			
kernel calls	intpts	azimuth	dip	intpts	azimuth	dip	diff. z
1	1	0/90	0/90	1	0/90	0/90	1
10	1	0/90	0/90	1	0/90	0/90	10
250	5	0/90	0/90	5	0/90	0/90	10
500	5	arb.	0/90	5	0/90	0/90	10
1000	5	arb.	0/90	5	arb.	0/90	10
1500	5	arb.	0/90	5	arb.	arb.	10
2250	5	arb.	arb.	5	arb.	arb.	10

Table 1: Example as a table for comparison: 1 source, 10 receiver (one or many frequencies).

## **Parallelisation**

If opt = 'parallel', six (\*) of the most time-consuming statements are calculated by using the numexpr package (https://github.com/pydata/numexpr/wiki/Numexpr-Users-Guide). These statements are all in the kernel-functions greenfct, reflections, and fields, and all involve  $\Gamma$  in one way or another, often calculating square roots or exponentials. As  $\Gamma$  has dimensions (#frequencies, #offsets, #layers, #lambdas), it can become fairly big.

The package numexpr has to be built with Intel's VML, otherwise it won't be used. You can check if it uses VML with

```
>>> import numexpr
>>> numexpr.use_vml
```

The module numexpr uses by default all available cores up to a maximum of 8. You can change this behaviour to a lower or a higher value with the following command (in the example it is changed to 4):

>>> import numexpr
>>> numexpr.set\_num\_threads(4)

This parallelisation will make empymod faster (by using more threads) if you calculate a lot of offsets/frequencies at once, but slower for few offsets/frequencies. Best practice is to check first which one is faster. (You can use the benchmark-notebook in the empymod/empymod-examples-repository.)

(\*) These statements are (following the notation of [HuTS15]):  $\Gamma$  (below eq. 19);  $W_n^{u,d}$  (eq. 74),  $r_n^{\pm}$  (eq. 65);  $R_n^{\pm}$  (eq. 64);  $P_s^{u,d;\pm}$  (eq. 81);  $M_s$  (eq. 82), and their corresponding bar-ed versions provided in the appendix (e.g.  $\overline{\Gamma}$ ). In big models, more than 95% of the calculation is spent in the calculation of these six equations, and most of the time therefore in np.sqrt and np.exp, or generally in numpy-ufuncs which are implemented and executed in compiled C-code. For smaller models or if transforms with interpolations are used then all the other parts also start to play a role. However, those models generally execute comparably fast.

## Lagged Convolution and Splined Transforms

Both Hankel and Fourier DLF have three options, which can be controlled via the htarg['pts\_per\_dec'] and ftarg['pts\_per\_dec'] parameters:

- pts\_per\_dec=0 : *Standard DLF*;
- pts\_per\_dec<0 : *Lagged Convolution DLF*: Spacing defined by filter base, interpolation is carried out in the input domain;
- pts\_per\_dec>0: *Splined DLF*: Spacing defined by pts\_per\_dec, interpolation is carried out in the output domain.

Similarly, interpolation can be used for QWE by setting pts\_per\_dec to a value bigger than 0.

The Lagged Convolution and Splined options should be used with caution, as they use interpolation and are therefore less precise than the standard version. However, they can significantly speed up *QWE*, and massively speed up *DLF*. Additionally, the interpolated versions minimizes memory requirements a lot. Speed-up is greater if all source-receiver angles are identical. Note that setting pts\_per\_dec to something else than 0 to calculate only one offset (Hankel) or only one time (Fourier) will be slower than using the standard version. Similarly, the standard version is usually the fastest when using the parallel option (numexpr).

QWE: Good speed-up is also achieved for QWE by setting maximt as low as possible. Also, the higher negative is, the higher the speed-up will be.

DLF: Big improvements are achieved for long DLF-filters and for many offsets/frequencies (thousands).

**Warning:** Keep in mind that setting pts\_per\_dec to something else than 0 uses interpolation, and is therefore not as accurate as the standard version. Use with caution and always compare with the standard version to verify if you can apply interpolation to your problem at hand!

Be aware that *QUAD* (Hankel transform) *always* use the splined version and *always* loops over offsets. The Fourier transforms *FFTlog*, *QWE*, and *FFT* always use interpolation too, either in the frequency or in the time domain. With the *DLF* Fourier transform (sine and cosine transforms) you can choose between no interpolation and interpolation (splined or lagged).

The splined versions of QWE check whether the ratio of any two adjacent intervals is above a certain threshold (steep end of the wavenumber or frequency spectrum). If it is, it carries out QUAD for this interval instead of QWE. The threshold is stored in diff\_quad, which can be changed within the parameter htarg and ftarg.

For a graphical explanation of the differences between standard DLF, lagged convolution DLF, and splined DLF for the Hankel and the Fourier transforms see the notebook 7a\_DLF-Standard-Lagged-Splined in the empymod-examples repository.

## Looping

By default, you can calculate many offsets and many frequencies all in one go, vectorized (for the DLF), which is the default. The loop parameter gives you the possibility to force looping over frequencies or offsets. This parameter can have severe effects on both runtime and memory usage. Play around with this factor to find the fastest version for your problem at hand. It ALWAYS loops over frequencies if ht = 'QWE'/'QUAD' or if ht = 'FHT' and pts\_per\_dec!=0 (Lagged Convolution or Splined Hankel DLF). All vectorized is very fast if there are few offsets or few frequencies. If there are many offsets and many frequencies, looping over the smaller of the two will be faster. Choosing the right looping together with opt = 'parallel' can have a huge influence.

## Vertical components and xdirect

Calculating the direct field in the wavenumber-frequency domain (xdirect=False; the default) is generally faster than calculating it in the frequency-space domain (xdirect=True).

However, using xdirect = True can improve the result (if source and receiver are in the same layer) to calculate:

- the vertical electric field due to a vertical electric source,
- configurations that involve vertical magnetic components (source or receiver),
- all configurations when source and receiver depth are exactly the same.

The Hankel transforms methods are having sometimes difficulties transforming these functions.

## Time-domain land CSEM

The derivation, as it stands, has a near-singular behaviour in the wavenumber-frequency domain when  $\kappa^2 = \omega^2 \epsilon \mu$ . This can be a problem for land-domain CSEM calculations if source and receiver are located at the surface between air and subsurface. Because most transforms do not sample the wavenumber-frequency domain sufficiently to catch this near-singular behaviour (hence not smooth), which then creates noise at early times where the signal should be zero. To avoid the issue simply set epermH[0] = epermV[0] = 0, hence the relative electric permittivity of the air to zero. This trick obviously uses the diffusive approximation for the air-layer, it therefore will not work for very high frequencies (e.g., GPR calculations).

This trick works fine for all horizontal components, but not so much for the vertical component. But then it is not feasible to have a vertical source or receiver *exactly* at the surface. A few tips for these cases: The receiver can be put pretty close to the surface (a few millimeters), but the source has to be put down a meter or two, more for the case of vertical source AND receiver, less for vertical source OR receiver. The results are generally better if the source is put deeper than the receiver. In either case, the best is to first test the survey layout against the analytical result (using empymod.analytical with solution='dhs') for a half-space, and subsequently model more complex cases.

# 5.1.8 License

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# 5.2 Roadmap

A collection of ideas of what could be added or improved in empymod. Please get in touch if you would like to tackle one of these problems!

- Additional modelling routines
  - tdem (TEM) [empymod#8]: Issues that have to be addressed: ramp waveform, windowing, loop integration, zero-offset (coincident loop).
    - \* in-loop
    - \* coincident loop
    - \* loop-loop
    - \* arbitrary shaped loops
  - Ramp waveform [empymod#7]
  - Arbitrary waveform [empymod#7]
  - Improve the GPR-routine [empymod#9]
  - Load and save functions to easily store and load model information (resistivity model, acquisition parameters, and modelling parameters) together with the modelling data (using pickle or shelve). Probably easier after implementation of the abstraction [empymod#14].
- Inversion [empymod#20]: Inversion routines, preferably a selection of different ones.
  - Add some clever checks, e.g. as in Key (2012): abort loops if the field is strongly attenuated.
- Additional (semi-)analytical functions (where possible)
  - Complete full-space (electric and magnetic source and receiver); space-time domain
  - Extend diffusive half-space solution to magnetic sources and receivers; space-frequency and space-time domains
  - Complete half-space
- Transforms
  - Fourier
    - \* Change fft to use discrete sine/cosine transforms instead, as all other Fourier transforms
    - \* If previous step is successful, clean up the internal decisions (utils.check\_time) when to use sine/cosine transform (not consistent at the moment, some choice only exists with ffht impulse responses, fqwe and fftlog use sine for impulse, and all three use sine for step-on responses and cosine for step-off responses)
  - Hankel
    - $\ast\,$  Add the fht-module from FFTLog for the Hankel transform.
  - Hankel and Fourier
    - \* Include the method outlined by Mulder et al., 2008, Geophysics (piecewise-cubic Hermite interpolation with a FFT) to try to further speed-up the splined versions.
- Extend examples (example-notebooks)
  - Add different methods (e.g. DC)
  - Reproduce published results
- Abstraction of the code [empymod#14].
- Replace (and extend) the use of numexpr with numba.
- GUI.

• Move empymod from channel 'prisae' to 'conda-forge' (pros/cons?).

# 5.3 Changelog

# 5.3.1 v1.8.3 - 2019-07-05

- Use scooby for Versions (printinfo), change name to Report.
- DOC: Correct return statement if mrec=True.
- Typos and correct links for new asv/bench.
- Bump requirement to SciPy>=1.0.0, remove warning regarding memory leak in SciPy 0.19.0.

# 5.3.2 v1.8.2 - 2019-04-26

- pts\_per\_dec are now floats, not integers, which gives more flexibility.
- Bugfix: pts\_per\_dec for DLF was actually points per e, not per decade, as the natural logarithm was used.
- New Versions-class; improvement over the versions-function, as it automatically detects if it can print html or not.
- Maintenance: Update np.load in tests with allow\_pickle=True for changes in numpy v1.16.3.
- Lots of changes to accommodate emg3d within the empymod-org:
  - Adjust website, move stuff from website into README.md.
  - /empymod/example-notebooks -> /empymod/empymod-examples.
  - /empymod/asv -> /empymod/empymod-asv (and therefore now available at empymod.github.io/empymod-asv).
  - /empymod/bench -> /empymod/empymod-bench.
- Move manual from empymod/\_\_init\_\_.py to the docs/manual.rst, and the references to its own file. Change reference style.
- Move credits for initial funding from the license-section of the manual to CREDITS.rst, where it belongs.

# 5.3.3 v1.8.1 - *2018-11-20*

- Many little improvements in the documentation.
- Some code improvements through the use of codacy.
- Remove testing of Python 3.4; officially supported are now Python 3.5-3.7.
- Version of the filter article (DLF) in geophysics and of the CSEM book.

# 5.3.4 v1.8.0 - 2018-10-26

- model.bipole, model.dipole, and model.analytical have now a hook which users can exploit to insert their own calculation of etaH, etaV, zetaH, and zetaV. This can be used, for instance, to model a Cole-Cole IP survey. See the manual or the example-notebooks for more information.
- model.wavenumber renamed to model.dipole\_k to avoid name clash with kernel. wavenumber. For now model.wavenumber continues to exist, but raises a depreciation warning.
- xdirect default value changed from True to False.

• Possibility to provide interpolated points (int\_pts) to transform.dlf.

The following changes are backwards incompatible if you directly used transform.fht, transform.hqwe, or transform.hquad. Nothing changes for the user-facing routines in model:

- empymod.fem now passes factAng to empymod.transform, not angle; this saves some time if looped over offsets or frequencies, as it is not repeatedly calculated within empymod.transform.
- Use get\_spline\_values in empymod.fem for Hankel DLF, instead of in empymod.fht. Gives a speed-up if looped over offsets or frequencies. Should be in utils, but that would be heavily backwards incompatible. Move there in version 2.0.

# 5.3.5 v1.7.3 - 2018-07-16

- Small improvements related to speed as a result of the benchmarks introduced in v1.7.2:
  - Kernels which do not exist for a given ab are now returned as None from kernel.wavenumber instead of arrays of zeroes. This permits for some time saving in the transforms. This change is backwards incompatible if you directly used kernel.wavenumber. Nothing changes for the user-facing routines in model.
  - Adjustments in transform with regard to the None returned by kernel.wavenumber. The kernels are not checked anymore if they are all zeroes (which can be slow for big arrays). If they are not None, they will be processed.
  - Various small improvements for speed to transform.dlf (i.e. factAng; log10/log; rearranging).

# 5.3.6 v1.7.2 - 2018-07-07

- Benchmarks: empymod has now a benchmark suite, see empymod/asv.
- Fixed a bug in bipole for time-domain responses with several receivers or sources with different depths. (Simply failed, as wrong dimension was provided to tem).
- Small improvements:
  - Various simplifications or cleaning of the code base.
  - Small change (for speed) in check if kernels are empty in transform.dlf and transform.qwe.

# 5.3.7 v1.7.1 - 2018-06-19

- New routines in empymod.filters.DigitalFilter: Filters can now be saved to or loaded from pure ascii-files.
- Filters and inversion result from empymod.scripts.fdesign are now by default saved in plain text. The filters with their internal routine, the inversion result with np.savetxt. Compressed saving can be achieved by giving a name with a '.gz'-ending.
- Change in empymod.utils:
  - Renamed \_min\_param to \_min\_res.
  - Anisotropy aniso is no longer directly checked for its minimum value. Instead, res\*aniso\*\*2, hence vertical resistivity, is checked with \_min\_res, and anisotropy is subsequently re-calculated from it.
  - The parameters epermH, epermV, mpermH, and mpermV can now be set to 0 (or any positive value) and do not depend on \_min\_param.
- printinfo: Generally improved; prints now MKL-info (if available) independently of numexpr.
- Simplification of kernel.reflections through re-arranging.

- Bug fixes
- Version of re-submission of the DLF article to geophysics.

# 5.3.8 v1.7.0 - 2018-05-23

Merge empyscripts into empymod under empymod.scripts.

- Clear separation between mandatory and optional imports:
  - Mandatory:
    - \* numpy
    - \* scipy
  - Optional:
    - \* numexpr(for empymod.kernel)
    - \* matplotlib(for empymod.scripts.fdesign)
    - \* IPython (for empymod.scripts.printinfo)
- Broaden namespace of empymod. All public functions from the various modules and the modules from empymod.scripts are now available under empymod directly.

# 5.3.9 v1.6.2 - 2018-05-21

These changes should make calculations using QWE and QUAD for the Hankel transform for cases which do not require all kernels faster; sometimes as much as twice as fast. However, it might make calculations which do require all kernels a tad slower, as more checks had to be included. (Related to [empymod#11]; basically including for QWE and QUAD what was included for DLF in version 1.6.0.)

- transform:
  - dlf:
    - \* Improved by avoiding unnecessary multiplications/summations for empty kernels and applying the angle factor only if it is not 1.
    - \* Empty/unused kernels can now be input as None, e.g. signal=(PJ0, None, None).
    - \* factAng is new optional for the Hankel transform, as is ab.
  - hqwe: Avoids unnecessary calculations for zero kernels, improving speed for these cases.
  - hquad, quad: Avoids unnecessary calculations for zero kernels, improving speed for these cases.
- kernel:
  - Simplify wavenumber
  - Simplify angle\_factor

# 5.3.10 v1.6.1 - 2018-05-05

Secondary field calculation.

- Add the possibility to calculate secondary fields only (excluding the direct field) by passing the argument xdirect=None. The complete xdirect-signature is now (only affects calculation if src and rec are in the same layer):
  - If True, direct field is calculated analytically in the frequency domain.
  - If False, direct field is calculated in the wavenumber domain.

- If None, direct field is excluded from the calculation, and only reflected fields are returned (secondary field).
- Bugfix in model.analytical for ab=[36, 63] (zeroes) [empymod#16].

# 5.3.11 v1.6.0 - 2018-05-01

This release is not completely backwards compatible for the main modelling routines in empymod.model, but almost. Read below to see which functions are affected.

- Improved Hankel DLF [empymod#11]. empymod.kernel.wavenumber always returns three kernels, PJ0, PJ1, and PJ0b. The first one is angle-independent, the latter two depend on the angle. Now, depending of what source-receiver configuration is chosen, some of these might be zero. If-statements were now included to avoid the calculation of the DLF, interpolation, and reshaping for 0-kernels, which improves speed for these cases.
- Unified DLF arguments [empymod#10].

These changes are backwards compatible for all main modelling routines in empymod.model. However, they are not backwards compatible for the following routines:

- empymod.model.fem (removed use\_spline),
- empymod.transform.fht (removed use\_spline),
- empymod.transform.hqwe (removed use\_spline),
- empymod.transform.quad(removeduse\_spline),
- empymod.transform.dlf(lagged, splined => pts\_per\_dec),
- empymod.utils.check\_opt (no longer returns use\_spline),
- empymod.utils.check\_hankel (changes in pts\_per\_dec), and
- empymod.utils.check\_time (changes in pts\_per\_dec).

The function empymod.utils.spline\_backwards\_hankel can be used for backwards compatibility.

Now the Hankel and Fourier DLF have the same behaviour for pts\_per\_dec:

- pts\_per\_dec = 0: Standard DLF,
- pts\_per\_dec < 0: Lagged Convolution DLF, and
- pts\_per\_dec > 0: Splined DLF.

There is one exception which is not backwards compatible: Before, if opt=None and htarg={pts\_per\_dec: != 0}, the pts\_per\_dec was not used for the FHT and the QWE. New, this will be used according to the above definitions.

• Bugfix in model.wavenumber for ab=[36, 63] (zeroes).

# 5.3.12 v1.5.2 - 2018-04-25

- DLF improvements:
  - Digital linear filter (DLF) method for the Fourier transform can now be carried out without spline, providing 0 for pts\_per\_dec (or any integer smaller than 1).
  - Combine kernel from fht and ffht into dlf, hence separate DLF from other calculations, as is done with QWE (qwe for hqwe and fqwe).
  - Bug fix regarding transform.get\_spline\_values; a DLF with pts\_per\_dec can now be shorter then the corresponding filter.

# 5.3.13 v1.5.1 - 2018-02-24

- Documentation:
  - Simplifications: avoid duplication as much as possible between the website (empymod.github.io), the manual (empymod.readthedocs.io), and the README (github.com/empymod/empymod).
    - \* Website has now only *Features* and *Installation* in full, all other information comes in the form of links.
    - \* README has only information in the form of links.
    - \* Manual contains the README, and is basically the main document for all information.
  - Improvements: Change some remaining md-syntax to rst-syntax.
  - FHT -> DLF: replace FHT as much as possible, without breaking backwards compatibility.

# 5.3.14 v1.5.0 - 2018-01-02

- Minimum parameter values can now be set and verified with utils.set\_minimum and utils. get\_minimum.
- New Hankel filter wer\_201\_2018.
- opt=parallel has no effect if numexpr is not built against Intel's VML. (Use import numexpr; numexpr.use\_vml to see if your numexpr uses VML.)
- Bug fixes
- Version of manuscript submission to geophysics for the DLF article.

# 5.3.15 v1.4.4 - 2017-09-18

[This was meant to be 1.4.3, but due to a setup/pypi/anaconda-issue I had to push it to 1.4.4; so there isn't really a version 1.4.3.]

- Add TE/TM split to diffusive ee-halfspace solution.
- Improve kernel.wavenumber for fullspaces.
- Extended fQWE and fftlog to be able to use the cosine-transform. Now the cosine-transform with the real-part frequency response is used internally if a switch-off response (signal=-1) is required, rather than calculating the switch-on response (with sine-transform and imaginary-part frequency response) and subtracting it from the DC value.
- Bug fixes

# 5.3.16 v1.4.2 - 2017-06-04

- Bugfix: Fixed squeeze in model.analytical with solution='dsplit'.
- Version of final submission of manuscript to Geophysics.

# 5.3.17 v1.4.1 - 2017-05-30

[This was meant to be 1.4.0, but due to a setup/pypi/anaconda-issue I had to push it to 1.4.1; so there isn't really a version 1.4.0.]

- New home: empymod.github.io as entry point, and the project page on github.com/empymod. All empymod-repos moved to the new home.
  - /prisae/empymod -> /empymod/empymod

- /prisae/empymod-notebooks -> /empymod/example-notebooks
- /prisae/empymod-geo2017 -> /empymod/article-geo2017
- /prisae/empymod-tle2017 -> /empymod/article-tle2017
- Modelling routines:
  - New modelling routine model.analytical, which serves as a front-end to kernel. fullspace or kernel.halfspace.
  - Remove legacy routines model.time and model.frequency. They are covered perfectly by model.dipole.
  - Improved switch-off response (calculate and subtract from DC).
  - xdirect adjustments:
    - \* isfullspace now respects xdirect.
    - \* Removed xdirect from model.wavenumber (set to False).
- Kernel:
  - Modify kernel.halfspace to use same input as other kernel functions.
  - Include time-domain ee halfspace solution into kernel.halfspace; possible to obtain direct, reflected, and airwave separately, as well as only fullspace solution (all for the diffusive approximation).

## 5.3.18 v1.3.0 - 2017-03-30

- Add additional transforms and improve QWE:
  - Conventional adaptive quadrature (QUADPACK) for the Hankel transform;
  - Conventional FFT for the Fourier transform.
  - Add diff\_quad to htarg/ftarg of QWE, a switch parameter for QWE/QUAD.
  - Change QWE/QUAD switch from comparing first interval to comparing all intervals.
  - Add parameters for QUAD (a, b, limit) into htarg/ftarg for QWE.
- Allow htarg/ftarg as dict additionally to list/tuple.
- Improve model.gpr.
- Internal changes:
  - Rename internally the sine/cosine filter from fft to ffht, because of the addition of the Fast Fourier Transform fft.
- · Clean-up repository
  - Move notebooks to /prisae/empymod-notebooks
  - Move publications/Geophysics2017 to /prisae/empymod-geo2017
  - Move publications/TheLeadingEdge2017 to /prisae/empymod-tle2017
- Bug fixes and documentation improvements

# 5.3.19 v1.2.1 - 2017-03-11

- Change default filter from key\_401\_2009 to key\_201\_2009 (because of warning regarding 401 pt filter in source code of DIPOLE1D.)
- Since 06/02/2017 installable via pip/conda.
- Bug fixes

# 5.3.20 v1.2.0 - 2017-02-02

- New routine:
  - General modelling routine bipole (replaces srcbipole): Model the EM field for arbitrarily oriented, finite length bipole sources and receivers.
- Added a test suite:
  - Unit-tests of small functions.
  - Framework-tests of the bigger functions:
    - \* Comparing to status quo (regression tests),
    - \* Comparing to known analytical solutions,
    - \* Comparing different options to each other,
    - \* Comparing to other 1D modellers (EMmod, DIPOLE1D, GREEN3D).
  - Incorporated with Travis CI and Coveralls.
- Internal changes:
  - Add kernel count (printed if verb > 1).
  - numexpr is now only required if opt=='parallel'. If numexpr is not found, opt is reset to None and a warning is printed.
  - Cleaned-up wavenumber-domain routine.
  - theta/phi -> azimuth/dip; easier to understand.
  - Refined verbosity levels.
  - Lots of changes in utils, with regards to the new routine bipole and with regards to verbosity. Moved all warnings out from transform and model into utils.
- Bug fixes

# 5.3.21 v1.1.0 - *2016-12-22*

- New routines:
  - New srcbipole modelling routine: Model an arbitrarily oriented, finite length bipole source.
  - Merge frequency and time into dipole. (frequency and time are still available.)
  - dipole now supports multiple sources.
- Internal changes:
  - Replace get\_Gauss\_Weights with scipy.special.p\_roots
  - jv(0,x), jv(1,x) -> j0(x), j1(x)
  - Replace param\_shape in utils with \_check\_var and \_check\_shape.
  - Replace xco and yco by angle in kernel.fullspace
  - Replace fftlog with python version.
  - Additional sine-/cosine-filters: key\_81\_CosSin\_2009, key\_241\_CosSin\_2009, and key\_601\_CosSin\_2009.
- Bug fixes

## 5.3.22 v1.0.0 - 2016-11-29

• Initial release; state of manuscript submission to geophysics.

# 5.4 Credits

Thanks to

- Jürg Hunziker, Kerry Key, and Evert Slob for answering all my questions regarding their codes and publications (Hunziker et al., 2015, Key, 2009, Key, 2012, Slob et al., 2010).
- Evert Slob for the feedback and interaction during the creation of the add-on tmtemod, which was developed for the creation of github.com/empymod/csem-ziolkowski-and-slob.
- Kerry Key and Evert Slob for their inputs and feedback during the development of the add-on fdesign (see github.com/empymod/article-fdesign).

**Note:** This software was initially (till 01/2017) developed with funding from *The Mexican National Council of Science and Technology (Consejo Nacional de Ciencia y Tecnología*, http://www.conacyt.gob.mx), carried out at *The Mexican Institute of Petroleum IMP (Instituto Mexicano del Petróleo*, http://www.gob.mx/imp).

# 5.5 Code

## 5.5.1 model - Model EM-responses

EM-modelling routines. The implemented routines might not be the fastest solution to your specific problem. Use these routines as template to create your own, problem-specific modelling routine!

### **Principal routines:**

- bipole
  - dipole

The main routine is bipole, which can model bipole source(s) and bipole receiver(s) of arbitrary direction, for electric or magnetic sources and receivers, both in frequency and in time. A subset of bipole is dipole, which models infinitesimal small dipoles along the principal axes x, y, and z.

Further routines are:

- analytical: Calculate analytical fullspace and halfspace solutions.
- dipole\_k: Calculate the electromagnetic wavenumber-domain solution.
- gpr: Calculate the Ground-Penetrating Radar (GPR) response.

The dipole\_k routine can be used if you are interested in the wavenumber-domain result, without Hankel nor Fourier transform. It calls straight the kernel. The gpr-routine convolves the frequency-domain result with a wavelet, and applies a gain to the time-domain result. This function is still experimental.

### The modelling routines make use of the following two core routines:

- fem: Calculate wavenumber-domain electromagnetic field and carry out the Hankel transform to the frequency domain.
- tem: Carry out the Fourier transform to time domain after fem.

Return the electromagnetic field due to an electromagnetic source.

Calculate the electromagnetic frequency- or time-domain field due to arbitrary finite electric or magnetic bipole sources, measured by arbitrary finite electric or magnetic bipole receivers. By default, the electromagnetic response is normalized to to source and receiver of 1 m length, and source strength of 1 A.

#### Parameters

src, rec [list of floats or arrays]

#### Source and receiver coordinates (m):

- [x0, x1, y0, y1, z0, z1] (bipole of finite length)
- [x, y, z, azimuth, dip] (dipole, infinitesimal small)

#### **Dimensions:**

- The coordinates x, y, and z (dipole) or x0, x1, y0, y1, z0, and z1 (bipole) can be single values or arrays.
- The variables x and y (dipole) or x0, x1, y0, and y1 (bipole) must have the same dimensions.
- The variable z (dipole) or z0 and z1 (bipole) must either be single values or having the same dimension as the other coordinates.
- The variables azimuth and dip must be single values. If they have different angles, you have to use the bipole-method (with srcpts/recpts = 1, so it is calculated as dipoles).

Angles (coordinate system is left-handed, positive z down (East-North-Depth):

- azimuth (°): horizontal deviation from x-axis, anti-clockwise.
- dip (°): vertical deviation from xy-plane downwards.

Sources or receivers placed on a layer interface are considered in the upper layer.

**depth** [list] Absolute layer interfaces z (m); #depth = #res - 1 (excluding +/- infinity).

**res** [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1.

Alternatively, res can be a dictionary. See the main manual of empymod too see how to exploit this hook to re-calculate etaH, etaV, zetaH, and zetaV, which can be used to, for instance, use the Cole-Cole model for IP.

**freqtime** [array\_like] Frequencies f (Hz) if signal == None, else times t (s); (f, t > 0).

**signal** [{None, 0, 1, -1}, optional]

#### Source signal, default is None:

- None: Frequency-domain response
- -1 : Switch-off time-domain response
- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response
- **aniso** [array\_like, optional] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res. Defaults to ones.
- **epermH, epermV** [array\_like, optional] Relative horizontal/vertical electric permittivities epsilon\_h/epsilon\_v (-); #epermH = #epermV = #res. Default is ones.

- **mpermH, mpermV** [array\_like, optional] Relative horizontal/vertical magnetic permeabilities mu\_h/mu\_v (-); #mpermH = #mpermV = #res. Default is ones.
- **msrc, mrec** [boolean, optional] If True, source/receiver (msrc/mrec) is magnetic, else electric. Default is False.

srcpts, recpts [int, optional]

#### Number of integration points for bipole source/receiver, default is 1:

- srcpts/recpts < 3 : bipole, but calculated as dipole at centre
- srcpts/recpts >= 3 : bipole

#### strength [float, optional]

### Source strength (A):

- If 0, output is normalized to source and receiver of 1 m length, and source strength of 1 A.
- If != 0, output is returned for given source and receiver length, and source strength.

Default is 0.

### xdirect [bool or None, optional]

### Direct field calculation (only if src and rec are in the same layer):

- If True, direct field is calculated analytically in the frequency domain.
- If False, direct field is calculated in the wavenumber domain.
- If None, direct field is excluded from the calculation, and only reflected fields are returned (secondary field).

Defaults to False.

**ht** [{'fht', 'qwe', 'quad'}, optional] Flag to choose either the *Digital Linear Filter* method (FHT, *Fast Hankel Transform*), the *Quadrature-With-Extrapolation* (QWE), or a simple *Quadrature* (QUAD) for the Hankel transform. Defaults to 'fht'.

htarg [dict or list, optional]

#### Depends on the value for ht:

- If ht = 'fht': [fhtfilt, pts\_per\_dec]:
  - fhtfilt: string of filter name in empymod.filters or the filter method itself. (default: empymod.filters.key\_201\_2009())
  - pts\_per\_dec: points per decade; (default: 0)
    - \* If 0: Standard DLF.
    - \* If < 0: Lagged Convolution DLF.
    - \* If > 0: Splined DLF
- If ht = 'qwe': [rtol, atol, nquad, maxint, pts\_per\_dec,

diff\_quad, a, b, limit]:

- rtol: relative tolerance (default: 1e-12)
- atol: absolute tolerance (default: 1e-30)
- nquad: order of Gaussian quadrature (default: 51)
- maxint: maximum number of partial integral intervals (default: 40)
- pts\_per\_dec: points per decade; (default: 0)
  - \* If 0, no interpolation is used.

- \* If > 0, interpolation is used.
- diff\_quad: criteria when to swap to QUAD (only relevant if opt='spline') (default: 100)
- a: lower limit for QUAD (default: first interval from QWE)
- b: upper limit for QUAD (default: last interval from QWE)
- limit: limit for quad (default: maxint)
- If ht = 'quad': [atol, rtol, limit, lmin, lmax, pts\_per\_dec]:
  - rtol: relative tolerance (default: 1e-12)
  - atol: absolute tolerance (default: 1e-20)
  - limit: An upper bound on the number of subintervals used in the adaptive algorithm (default: 500)
  - Imin: Minimum wavenumber (default 1e-6)
  - lmax: Maximum wavenumber (default 0.1)
  - pts\_per\_dec: points per decade (default: 40)

The values can be provided as dict with the keywords, or as list. However, if provided as list, you have to follow the order given above. A few examples, assuming ht = qwe:

- Only changing rtol: { 'rtol': 1e-4} or [1e-4] or 1e-4
- Changing rtol and nquad: { 'rtol': 1e-4, 'nquad': 101 } or [1e-4, '', 101]
- **Only changing diff\_quad:** {'diffquad': 10} or ['', '', '', '', 10]
- ft [{'sin', 'cos', 'qwe', 'fftlog', 'fft'}, optional] Only used if signal != None. Flag to choose either the Digital Linear Filter method (Sine- or Cosine-Filter), the Quadrature-With-Extrapolation (QWE), the FFTLog, or the FFT for the Fourier transform. Defaults to 'sin'.

ftarg [dict or list, optional]

### Only used if signal != None. Depends on the value for ft:

- If ft = 'sin' or 'cos': [fftfilt, pts\_per\_dec]:
  - fftfilt: string of filter name in empymod.filters or the filter method itself. (Default: empymod.filters. key\_201\_CosSin\_2012())
  - pts\_per\_dec: points per decade; (default: -1)
    - \* If 0: Standard DLF.
    - \* If < 0: Lagged Convolution DLF.
    - \* If > 0: Splined DLF
- If ft = 'qwe': [rtol, atol, nquad, maxint, pts\_per\_dec]:
  - rtol: relative tolerance (default: 1e-8)
  - atol: absolute tolerance (default: 1e-20)
  - nquad: order of Gaussian quadrature (default: 21)
  - maxint: maximum number of partial integral intervals (default: 200)
  - pts\_per\_dec: points per decade (default: 20)
  - diff\_quad: criteria when to swap to QUAD (default: 100)
  - a: lower limit for QUAD (default: first interval from QWE)

- b: upper limit for QUAD (default: last interval from QWE)
- limit: limit for quad (default: maxint)
- If ft = 'fftlog': [pts\_per\_dec, add\_dec, q]:
  - pts\_per\_dec: sampels per decade (default: 10)
  - add\_dec: additional decades [left, right] (default: [-2, 1])
  - q: exponent of power law bias (default: 0);  $-1 \le q \le 1$
- If ft = 'fft': [dfreq, nfreq, ntot]:
  - dfreq: Linear step-size of frequencies (default: 0.002)
  - nfreq: Number of frequencies (default: 2048)
  - ntot: Total number for FFT; difference between nfreq and ntot is padded with zeroes. This number is ideally a power of 2, e.g. 2048 or 4096 (default: nfreq).
  - pts\_per\_dec : points per decade (default: None)

Padding can sometimes improve the result, not always. The default samples from 0.002 Hz - 4.096 Hz. If pts\_per\_dec is set to an integer, calculated frequencies are logarithmically spaced with the given number per decade, and then interpolated to yield the required frequencies for the FFT.

The values can be provided as dict with the keywords, or as list. However, if provided as list, you have to follow the order given above. See htarg for a few examples.

**opt** [{None, 'parallel'}, optional]

#### **Optimization flag. Defaults to None:**

- None: Normal case, no parallelization nor interpolation is used.
- If 'parallel', the package numexpr is used to evaluate the most expensive statements. Always check if it actually improves performance for a specific problem. It can speed up the calculation for big arrays, but will most likely be slower for small arrays. It will use all available cores for these specific statements, which all contain Gamma in one way or another, which has dimensions (#frequencies, #offsets, #layers, #lambdas), therefore can grow pretty big. The module numexpr uses by default all available cores up to a maximum of 8. You can change this behaviour to your desired number of threads nthreads with numexpr.set\_num\_threads (nthreads).
- The value 'spline' is deprecated and will be removed. See htarg instead for the interpolated versions.

The option 'parallel' only affects speed and memory usage, whereas 'spline' also affects precision! Please read the note in the *README* documentation for more information.

loop [{None, 'freq', 'off'}, optional] Define if to calculate everything vectorized or if to loop over frequencies ('freq') or over offsets ('off'), default is None. It always loops over frequencies if ht = 'qwe' or if opt = 'spline'. Calculating everything vectorized is fast for few offsets OR for few frequencies. However, if you calculate many frequencies for many offsets, it might be faster to loop over frequencies. Only comparing the different versions will yield the answer for your specific problem at hand!

## **verb** [{0, 1, 2, 3, 4}, optional]

## Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.

- 2: Print additional runtime and kernel calls
- 3: Print additional start/stop, condensed parameter information.
- 4: Print additional full parameter information

#### Returns

EM [ndarray, (nfreq, nrec, nsrc)]

### Frequency- or time-domain EM field (depending on signal):

- If rec is electric, returns E [V/m].
- If rec is magnetic, returns H [A/m].

However, source and receiver are normalised (unless strength != 0). So for instance in the electric case the source strength is 1 A and its length is 1 m. So the electric field could also be written as [V/(A.m2)].

In the magnetic case the source strength is given by  $i\omega\mu_0 AI^e$ , where A is the loop area (m2), and  $I^e$  the electric source strength. For the normalized magnetic source  $A = 1m^2$  and  $I^e = 1Ampere$ . A magnetic source is therefore frequency dependent.

The shape of EM is (nfreq, nrec, nsrc). However, single dimensions are removed.

### See also:

*fem* Electromagnetic frequency-domain response. *tem* Electromagnetic time-domain response.

#### **Examples**

```
>>> import numpy as np
>>> from empymod import bipole
>>> # x-directed bipole source: x0, x1, y0, y1, z0, z1
>>> src = [-50, 50, 0, 0, 100, 100]
>>> # x-directed dipole source-array: x, y, z, azimuth, dip
>>> rec = [np.arange(1, 11)*500, np.zeros(10), 200, 0, 0]
>>> # layer boundaries
>>> depth = [0, 300, 1000, 1050]
>>> # layer resistivities
>>> res = [1e20, .3, 1, 50, 1]
>>> # Frequency
>>> freq = 1
>>> # Calculate electric field due to an electric source at 1 Hz.
>>> # [msrc = mrec = True (default)]
>>> EMfield = bipole(src, rec, depth, res, freq, verb=4)
:: empymod START ::
           [m] : 0 300 1000 1050
  depth
  aniso [-] : 1 1 1 1 1
epermH [-1 · ·
           [Ohm.m] : 1E+20 0.3 1 50 1
  res
  epermV
mpermH
mpermV
              [-]: 11111
              [-]: 11111
              [-]: 11111
   frequency [Hz] : 1
  Hankel : DLF (Fast Hankel Transform)
> Filter : Key 201 (2009)
    > DLF type : Standard
  Kernel Opt.
                   : None
  Loop over : None (all vectorized)
Source(s) : 1 bipole(s)
> intpts : 1 (as dipole)
```

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```
> length [m] : 100
    > х с
            [m] : O
             [m] : O
    > y_c
    > z_c [m] : 100
    > azimuth [°] : 0
    > dip [°] : 0
  Receiver(s) : 10 dipole(s)
    > x [m] : 500 - 5000 : 10 [min-max; #]
                 : 500 1000 1500 2000 2500 3000 3500 4000 4500 5000
    > y
              [m] :
                    0 - 0 : 10 [min-max; #]
                    0 0 0 0 0 0 0 0 0 0
                  :
    > z
              [m] :
                    200
    > azimuth [°] :
                    0
           [°] :
    > dip
                    0
                 : 11
  Required ab's
:: empymod END; runtime = 0:00:00.005536 :: 1 kernel call(s)
>>> print (EMfield)
 1.68809346e-10 -3.08303130e-10j -8.77189179e-12 -3.76920235e-11j
Γ
 -3.46654704e-12 -4.87133683e-12j -3.60159726e-13 -1.12434417e-12j
  1.87807271e-13 -6.21669759e-13j 1.97200208e-13 -4.38210489e-13j
  1.44134842e-13 -3.17505260e-13j 9.92770406e-14 -2.33950871e-13j
  6.75287598e-14 -1.74922886e-13j 4.62724887e-14 -1.32266600e-13j]
```

Return the electromagnetic field due to a dipole source.

Calculate the electromagnetic frequency- or time-domain field due to infinitesimal small electric or magnetic dipole source(s), measured by infinitesimal small electric or magnetic dipole receiver(s); sources and receivers are directed along the principal directions x, y, or z, and all sources are at the same depth, as well as all receivers are at the same depth.

Use the functions bipole to calculate dipoles with arbitrary angles or bipoles of finite length and arbitrary angle.

The function dipole could be replaced by bipole (all there is to do is translate ab into msrc, mrec, azimuth's and dip's). However, dipole is kept separately to serve as an example of a simple modelling routine that can serve as a template.

#### **Parameters**

**src, rec** [list of floats or arrays] Source and receiver coordinates (m): [x, y, z]. The xand y-coordinates can be arrays, z is a single value. The x- and y-coordinates must have the same dimension.

Sources or receivers placed on a layer interface are considered in the upper layer. **depth** [list] Absolute layer interfaces z (m); #depth = #res - 1 (excluding +/- infinity). **res** [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1.

Alternatively, res can be a dictionary. See the main manual of empymod too see how to exploit this hook to re-calculate etaH, etaV, zetaH, and zetaV, which can be used to, for instance, use the Cole-Cole model for IP.

freqtime [array\_like] Frequencies f (Hz) if signal == None, else times t (s); (f, t > 0).
signal [{None, 0, 1, -1}, optional]

#### Source signal, default is None:

- None: Frequency-domain response
- -1 : Switch-off time-domain response

- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response

**ab** [int, optional] Source-receiver configuration, defaults to 11.

		electric source			magnetic source		
		X	У	Z	X	У	Z
electric	X	11	12	13	14	15	16
receiver	У	21	22	23	24	25	26
	Z	31	32	33	34	35	36
magnetic	X	41	42	43	44	45	46
receiver	У	51	52	53	54	55	56
	Z	61	62	63	64	65	66

**aniso** [array\_like, optional] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res. Defaults to ones.

**epermH, epermV** [array\_like, optional] Relative horizontal/vertical electric permittivities epsilon\_h/epsilon\_v (-); #epermH = #epermV = #res. Default is ones.

**mpermH, mpermV** [array\_like, optional] Relative horizontal/vertical magnetic permeabilities mu\_h/mu\_v (-); #mpermH = #mpermV = #res. Default is ones.

xdirect [bool or None, optional]

### Direct field calculation (only if src and rec are in the same layer):

- If True, direct field is calculated analytically in the frequency domain.
- If False, direct field is calculated in the wavenumber domain.
- If None, direct field is excluded from the calculation, and only reflected fields are returned (secondary field).

Defaults to False.

**ht** [{'fht', 'qwe', 'quad'}, optional] Flag to choose either the *Digital Linear Filter* method (FHT, *Fast Hankel Transform*), the *Quadrature-With-Extrapolation* (QWE), or a simple *Quadrature* (QUAD) for the Hankel transform. Defaults to 'fht'.

htarg [dict or list, optional]

#### Depends on the value for ht:

- If ht = 'fht': [fhtfilt, pts\_per\_dec]:
  - fhtfilt: string of filter name in empymod.filters or the filter method itself. (default: empymod.filters. key\_201\_2009())
  - pts\_per\_dec: points per decade; (default: 0)
    - \* If 0: Standard DLF.
    - \* If < 0: Lagged Convolution DLF.
    - \* If > 0: Splined DLF

#### • If ht = 'qwe': [rtol, atol, nquad, maxint, pts\_per\_dec,

diff\_quad, a, b, limit]:

- rtol: relative tolerance (default: 1e-12)
- atol: absolute tolerance (default: 1e-30)
- nquad: order of Gaussian quadrature (default: 51)
- maxint: maximum number of partial integral intervals (default: 40)

- pts\_per\_dec: points per decade; (default: 0)
  - \* If 0, no interpolation is used.
  - \* If > 0, interpolation is used.
- diff\_quad: criteria when to swap to QUAD (only relevant if opt='spline') (default: 100)
- a: lower limit for QUAD (default: first interval from QWE)
- b: upper limit for QUAD (default: last interval from QWE)
- limit: limit for quad (default: maxint)
- If ht = 'quad': [atol, rtol, limit, lmin, lmax, pts\_per\_dec]:
  - rtol: relative tolerance (default: 1e-12)
  - atol: absolute tolerance (default: 1e-20)
  - limit: An upper bound on the number of subintervals used in the adaptive algorithm (default: 500)
  - Imin: Minimum wavenumber (default 1e-6)
  - lmax: Maximum wavenumber (default 0.1)
  - pts\_per\_dec: points per decade (default: 40)

The values can be provided as dict with the keywords, or as list. However, if provided as list, you have to follow the order given above. A few examples, assuming ht = qwe:

- Only changing rtol: { 'rtol': 1e-4} or [1e-4] or 1e-4
- Changing rtol and nquad: { 'rtol': 1e-4, 'nquad': 101 } or [1e-4, '', 101]
- **Only changing diff\_quad:** {'diffquad': 10} or ['', '', '', '', 10]
- ft [{'sin', 'cos', 'qwe', 'fftlog', 'fft'}, optional] Only used if signal != None. Flag to choose either the Digital Linear Filter method (Sine- or Cosine-Filter), the Quadrature-With-Extrapolation (QWE), the FFTLog, or the FFT for the Fourier transform. Defaults to 'sin'.
- ftarg [dict or list, optional]

Only used if signal != None. Depends on the value for ft:

- If ft = 'sin' or 'cos': [fftfilt, pts\_per\_dec]:
  - fftfilt: string of filter name in empymod.filters or the filter method itself. (Default: empymod.filters. key\_201\_CosSin\_2012())
  - pts\_per\_dec: points per decade; (default: -1)
    - \* If 0: Standard DLF.
    - \* If < 0: Lagged Convolution DLF.
    - \* If > 0: Splined DLF
- If ft = 'qwe': [rtol, atol, nquad, maxint, pts\_per\_dec]:
  - rtol: relative tolerance (default: 1e-8)
  - atol: absolute tolerance (default: 1e-20)
  - nquad: order of Gaussian quadrature (default: 21)
  - maxint: maximum number of partial integral intervals (default: 200)
  - pts\_per\_dec: points per decade (default: 20)

- diff\_quad: criteria when to swap to QUAD (default: 100)
- a: lower limit for QUAD (default: first interval from QWE)
- b: upper limit for QUAD (default: last interval from QWE)
- limit: limit for quad (default: maxint)
- If ft = 'fftlog': [pts\_per\_dec, add\_dec, q]:
  - pts\_per\_dec: sampels per decade (default: 10)
  - add\_dec: additional decades [left, right] (default: [-2, 1])
  - q: exponent of power law bias (default: 0);  $-1 \le q \le 1$
- If ft = 'fft': [dfreq, nfreq, ntot]:
  - dfreq: Linear step-size of frequencies (default: 0.002)
  - nfreq: Number of frequencies (default: 2048)
  - ntot: Total number for FFT; difference between nfreq and ntot is padded with zeroes. This number is ideally a power of 2, e.g. 2048 or 4096 (default: nfreq).
  - pts\_per\_dec : points per decade (default: None)

Padding can sometimes improve the result, not always. The default samples from 0.002 Hz - 4.096 Hz. If pts\_per\_dec is set to an integer, calculated frequencies are logarithmically spaced with the given number per decade, and then interpolated to yield the required frequencies for the FFT.

The values can be provided as dict with the keywords, or as list. However, if provided as list, you have to follow the order given above. See htarg for a few examples.

**opt** [{None, 'parallel'}, optional]

# **Optimization flag. Defaults to None:**

- None: Normal case, no parallelization nor interpolation is used.
- If 'parallel', the package numexpr is used to evaluate the most expensive statements. Always check if it actually improves performance for a specific problem. It can speed up the calculation for big arrays, but will most likely be slower for small arrays. It will use all available cores for these specific statements, which all contain Gamma in one way or another, which has dimensions (#frequencies, #offsets, #layers, #lambdas), therefore can grow pretty big. The module numexpr uses by default all available cores up to a maximum of 8. You can change this behaviour to your desired number of threads nthreads with numexpr.set\_num\_threads (nthreads).
- The value 'spline' is deprecated and will be removed. See htarg instead for the interpolated versions.

The option 'parallel' only affects speed and memory usage, whereas 'spline' also affects precision! Please read the note in the *README* documentation for more information.

- loop [{None, 'freq', 'off'}, optional] Define if to calculate everything vectorized or if to loop over frequencies ('freq') or over offsets ('off'), default is None. It always loops over frequencies if ht = 'qwe' or if opt = 'spline'. Calculating everything vectorized is fast for few offsets OR for few frequencies. However, if you calculate many frequencies for many offsets, it might be faster to loop over frequencies. Only comparing the different versions will yield the answer for your specific problem at hand!
- **verb** [{0, 1, 2, 3, 4}, optional]

#### Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.
- 2: Print additional runtime and kernel calls
- 3: Print additional start/stop, condensed parameter information.
- 4: Print additional full parameter information

# Returns

EM [ndarray, (nfreq, nrec, nsrc)]

#### Frequency- or time-domain EM field (depending on signal):

- If rec is electric, returns E [V/m].
- If rec is magnetic, returns H [A/m].

However, source and receiver are normalised. So for instance in the electric case the source strength is 1 A and its length is 1 m. So the electric field could also be written as [V/(A.m2)].

The shape of EM is (nfreq, nrec, nsrc). However, single dimensions are removed.

### See also:

bipole Electromagnetic field due to an electromagnetic source.fem Electromagnetic frequency-domain response.tem Electromagnetic time-domain response.

### **Examples**

```
>>> import numpy as np
>>> from empymod import dipole
>>> src = [0, 0, 100]
>>> rec = [np.arange(1, 11)*500, np.zeros(10), 200]
>>> depth = [0, 300, 1000, 1050]
>>> res = [1e20, .3, 1, 50, 1]
>>> EMfield = dipole(src, rec, depth, res, freqtime=1, verb=0)
>>> print(EMfield)
[ 1.68809346e-10 -3.08303130e-10j -8.77189179e-12 -3.76920235e-11j
-3.46654704e-12 -4.87133683e-12j -3.60159726e-13 -1.12434417e-12j
1.87807271e-13 -6.21669759e-13j 1.97200208e-13 -4.38210489e-13j
1.44134842e-13 -3.17505260e-13j 9.92770406e-14 -2.33950871e-13j
6.75287598e-14 -1.74922886e-13j 4.62724887e-14 -1.32266600e-13j]
```

Return the analytical full- or half-space solution.

Calculate the electromagnetic frequency- or time-domain field due to infinitesimal small electric or magnetic dipole source(s), measured by infinitesimal small electric or magnetic dipole receiver(s); sources and receivers are directed along the principal directions x, y, or z, and all sources are at the same depth, as well as all receivers are at the same depth.

In the case of a halfspace the air-interface is located at z = 0 m.

You can call the functions fullspace and halfspace in kernel.py directly. This interface is just to provide a consistent interface with the same input parameters as for instance for dipole.

This function yields the same result if solution='fs' as dipole, if the model is a fullspace. Included are:

• Full fullspace solution (solution='fs') for ee-, me-, em-, mm-fields, only frequency domain, [HuTS15].

- Diffusive fullspace solution (solution='dfs') for ee-fields, [SIHM10].
- Diffusive halfspace solution (solution='dhs') for ee-fields, [SIHM10].
- Diffusive direct- and reflected field and airwave (solution='dsplit') for ee-fields, [SIHM10].
- Diffusive direct- and reflected field and airwave (solution='dtetm') for ee-fields, split into TE and TM mode [SIHM10].

#### Parameters

- **src, rec** [list of floats or arrays] Source and receiver coordinates (m): [x, y, z]. The xand y-coordinates can be arrays, z is a single value. The x- and y-coordinates must have the same dimension.
- res [float] Horizontal resistivity rho\_h (Ohm.m).

Alternatively, res can be a dictionary. See the main manual of empymod too see how to exploit this hook to re-calculate etaH, etaV, zetaH, and zetaV, which can be used to, for instance, use the Cole-Cole model for IP.

freqtime [array\_like] Frequencies f (Hz) if signal == None, else times t (s); (f, t > 0).
solution [str, optional]

#### Defines which solution is returned:

- 'fs' : Full fullspace solution (ee-, me-, em-, mm-fields); f-domain.
- 'dfs' : Diffusive fullspace solution (ee-fields only).
- 'dhs' : Diffusive halfspace solution (ee-fields only).
- 'dsplit' [Diffusive direct- and reflected field and airwave] (ee-fields only).
- 'dtetm' [as dsplit, but direct fielt TE, TM; reflected field TE, TM,] and airwave (ee-fields only).

**signal** [{None, 0, 1, -1}, optional]

#### Source signal, default is None:

- None: Frequency-domain response
- -1 : Switch-off time-domain response
- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response

ab [int, optional] Source-receiver configuration, defaults to 11.

		electric source			magnetic source		
		X	У	Z	X	У	Z
electric	x	11	12	13	14	15	16
receiver	У	21	22	23	24	25	26
	Z	31	32	33	34	35	36
magnetic	x	41	42	43	44	45	46
receiver	У	51	52	53	54	55	56
	Z	61	62	63	64	65	66

aniso [float, optional] Anisotropy lambda = sqrt(rho\_v/rho\_h) (-); defaults to one.
 epermH, epermV [float, optional] Relative horizontal/vertical electric permittivity epsilon\_h/epsilon\_v (-); default is one. Ignored for the diffusive solution.

**mpermH, mpermV** [float, optional] Relative horizontal/vertical magnetic permeability mu\_h/mu\_v (-); default is one. Ignored for the diffusive solution.

**verb** [{0, 1, 2, 3, 4}, optional]

#### Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.

- 2: Print additional runtime
- 3: Print additional start/stop, condensed parameter information.
- 4: Print additional full parameter information

# Returns

**EM** [ndarray, (nfreq, nrec, nsrc)]

#### Frequency- or time-domain EM field (depending on signal):

- If rec is electric, returns E [V/m].
- If rec is magnetic, returns H [A/m].

However, source and receiver are normalised. So for instance in the electric case the source strength is 1 A and its length is 1 m. So the electric field could also be written as [V/(A.m2)].

The shape of EM is (nfreq, nrec, nsrc). However, single dimensions are removed.

If solution='dsplit', three ndarrays are returned: direct, reflect, air.

If solution='dtetm', five ndarrays are returned: direct\_TE, direct\_TM, re-flect\_TE, reflect\_TM, air.

## Examples

THIS FUNCTION IS EXPERIMENTAL, USE WITH CAUTION.

It is rather an example how you can calculate GPR responses; however, DO NOT RELY ON IT! It works only well with QUAD or QWE (quad, qwe) for the Hankel transform, and with FFT (fft) for the Fourier transform.

It calls internally dipole for the frequency-domain calculation. It subsequently convolves the response with a Ricker wavelet with central frequency cf. If signal!=None, it carries out the Fourier transform and applies a gain to the response.

For input parameters see the function dipole, except for:

#### Parameters

- cf [float] Centre frequency of GPR-signal, in Hz. Sensible values are between 10 MHz and 3000 MHz.
- **gain** [float] Power of gain function. If None, no gain is applied. Only used if signal!=None.

#### Returns

EM [ndarray] GPR response

Return the electromagnetic wavenumber-domain field.

Calculate the electromagnetic wavenumber-domain field due to infinitesimal small electric or magnetic dipole source(s), measured by infinitesimal small electric or magnetic dipole receiver(s); sources and receivers are directed along the principal directions x, y, or z, and all sources are at the same depth, as well as all receivers are at the same depth.

## Parameters

**src, rec** [list of floats or arrays] Source and receiver coordinates (m): [x, y, z]. The x- and y-coordinates can be arrays, z is a single value. The x- and y-coordinates must have the same dimension. The x- and y-coordinates only matter for the angle-dependent factor.

Sources or receivers placed on a layer interface are considered in the upper layer. **depth** [list] Absolute layer interfaces z (m); #depth = #res - 1 (excluding +/- infinity). **res** [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1. **freq** [array\_like] Frequencies f (Hz), used to calculate etaH/V and zetaH/V. **wavenumber** [array] Wavenumbers lambda (1/m)

**ab** [int, optional] Source-receiver configuration, defaults to 11.

		electric source			magnetic source		
		X	У	Z	X	У	Z
electric	x	11	12	13	14	15	16
receiver	У	21	22	23	24	25	26
	Z	31	32	33	34	35	36
magnetic	X	41	42	43	44	45	46
receiver	У	51	52	53	54	55	56
	Z	61	62	63	64	65	66

**aniso** [array\_like, optional] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res. Defaults to ones.

**epermH, epermV** [array\_like, optional] Relative horizontal/vertical electric permittivities epsilon\_h/epsilon\_v (-); #epermH = #epermV = #res. Default is ones.

**mpermH, mpermV** [array\_like, optional] Relative horizontal/vertical magnetic permeabilities mu\_h/mu\_v (-); #mpermH = #mpermV = #res. Default is ones.

**verb** [{0, 1, 2, 3, 4}, optional]

### Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.
- 2: Print additional runtime and kernel calls
- 3: Print additional start/stop, condensed parameter information.
- 4: Print additional full parameter information

#### Returns

PJ0, PJ1 [array]

#### Wavenumber-domain EM responses:

- PJ0: Wavenumber-domain solution for the kernel with a Bessel function of the first kind of order zero.
- PJ1: Wavenumber-domain solution for the kernel with a Bessel function of the first kind of order one.

### See also:

*dipole* Electromagnetic field due to an electromagnetic source (dipoles). *bipole* Electromagnetic field due to an electromagnetic source (bipoles).

*fem* Electromagnetic frequency-domain response. *tem* Electromagnetic time-domain response.

#### **Examples**

```
>>> import numpy as np
>>> from empymod.model import dipole k
>>> src = [0, 0, 100]
>>> rec = [5000, 0, 200]
>>> depth = [0, 300, 1000, 1050]
>>> res = [1e20, .3, 1, 50, 1]
>>> freq = 1
>>> wavenrs = np.logspace(-3.7, -3.6, 10)
>>> PJ0, PJ1 = dipole_k(src, rec, depth, res, freq, wavenrs, verb=0)
>>> print(PJ0)
[ -1.02638329e-08 +4.91531529e-09j -1.05289724e-08 +5.04222413e-09j
  -1.08009148e-08 +5.17238608e-09j -1.10798310e-08 +5.30588284e-09j
  -1.13658957e-08 +5.44279805e-09j -1.16592877e-08 +5.58321732e-09j
  -1.19601897e-08 +5.72722830e-09j -1.22687889e-08 +5.87492067e-09j
  -1.25852765e-08 +6.02638626e-09j -1.29098481e-08 +6.18171904e-09j]
>>> print (PJ1)
[ 1.79483705e-10 -6.59235332e-10j 1.88672497e-10 -6.93749344e-10j
  1.98325814e-10 -7.30068377e-10j 2.08466693e-10 -7.68286748e-10j
  2.19119282e-10 -8.08503709e-10j 2.30308887e-10 -8.50823701e-10j
  2.42062030e-10 -8.95356636e-10j 2.54406501e-10 -9.42218177e-10j
  2.67371420e-10 -9.91530051e-101 2.80987292e-10 -1.04342036e-091
```

This function is called from one of the above modelling routines. No input-check is carried out here. See the main description of model for information regarding input and output parameters.

This function can be directly used if you are sure the provided input is in the correct format. This is useful for inversion routines and similar, as it can speed-up the calculation by omitting input-checks.

empymod.model.tem(fEM, off, freq, time, signal, ft, ftarg, conv=True)

Return the time-domain response of the frequency-domain response fEM.

This function is called from one of the above modelling routines. No input-check is carried out here. See the main description of model for information regarding input and output parameters.

This function can be directly used if you are sure the provided input is in the correct format. This is useful for inversion routines and similar, as it can speed-up the calculation by omitting input-checks.

Depreciated. Use *dipole\_k* instead.

# 5.5.2 kernel – Kernel calculation

Kernel of empymod, calculates the wavenumber-domain electromagnetic response. Plus analytical full- and half-space solutions.

The functions wavenumber, angle\_factor, fullspace, greenfct, reflections, and fields are based on source files (specified in each function) from the source code distributed with [HuTS15], which can be found at software.seg.org/2015/0001. These functions are (c) 2015 by Hunziker et al. and the Society of Exploration Geophysicists, http://software.seg.org/disclaimer.txt. Please read the NOTICE-file in the root directory for more information regarding the involved licenses.

Calculate wavenumber domain solution.

Return the wavenumber domain solutions PJO, PJ1, and PJOb, which have to be transformed with a Hankel transform to the frequency domain. PJO/PJOb and PJ1 have to be transformed with Bessel functions of order  $0 (J_0)$  and  $1 (J_1)$ , respectively.

This function corresponds loosely to equations 105–107, 111–116, 119–121, and 123–128 in [HuTS15], and equally loosely to the file kxwmod.c.

[HuTS15] uses Bessel functions of orders 0, 1, and 2  $(J_0, J_1, J_2)$ . The implementations of the *Fast Hankel Transform* and the *Quadrature-with-Extrapolation* in transform are set-up with Bessel functions of order 0 and 1 only. This is achieved by applying the recurrence formula

$$J_2(kr) = \frac{2}{kr} J_1(kr) - J_0(kr) \; .$$

**Note:** PJ0 and PJ0b could theoretically be added here into one, and then be transformed in one go. However, PJ0b has to be multiplied by factAng later. This has to be done after the Hankel transform for methods which make use of spline interpolation, in order to work for offsets that are not in line with each other.

This function is called from one of the Hankel functions in transform. Consult the modelling routines in model for a description of the input and output parameters.

If you are solely interested in the wavenumber-domain solution you can call this function directly. However, you have to make sure all input arguments are correct, as no checks are carried out here.

```
empymod.kernel.angle_factor(angle, ab, msrc, mrec)
```

Return the angle-dependent factor.

The whole calculation in the wavenumber domain is only a function of the distance between the source and the receiver, it is independent of the angel. The angle-dependency is this factor, which can be applied to the corresponding parts in the wavenumber or in the frequency domain.

The angle\_factor corresponds to the sine and cosine-functions in Eqs 105-107, 111-116, 119-121, 123-128.

This function is called from one of the Hankel functions in transform. Consult the modelling routines in model for a description of the input and output parameters.

empymod.kernel.fullspace (*off, angle, zsrc, zrec, etaH, etaV, zetaH, zetaV, ab, msrc, mrec*) Analytical full-space solutions in the frequency domain.

$$\hat{G}^{ee}_{lphaeta},\hat{G}^{ee}_{3lpha},\hat{G}^{ee}_{33},\hat{G}^{em}_{lphaeta},\hat{G}^{em}_{lpha3}$$

This function corresponds to equations 45–50 in [HuTS15], and loosely to the corresponding files Gin11. F90, Gin12.F90, Gin13.F90, Gin22.F90, Gin23.F90, Gin31.F90, Gin32.F90, Gin33. F90, Gin41.F90, Gin42.F90, Gin43.F90, Gin51.F90, Gin52.F90, Gin53.F90, Gin61. F90, and Gin62.F90.

This function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

empymod.kernel.greenfct (zsrc, zrec, lsrc, lrec, depth, etaH, etaV, zetaH, zetaV, lambd, ab, xdirect,

*msrc*, *mrec*, *use\_ne\_eval*) Calculate Green's function for TM and TE.

iculate Green's function for TM and TE.

$$ilde{g}_{hh}^{tm}, ilde{g}_{hz}^{tm}, ilde{g}_{zh}^{tm}, ilde{g}_{zz}^{tm}, ilde{g}_{hh}^{te}, ilde{g}_{zz}^{te}$$

This function corresponds to equations 108–110, 117/118, 122; 89–94, A18–A23, B13–B15; 97–102 A26–A31, and B16–B18 in [HuTS15], and loosely to the corresponding files Gamma.F90, Wprop. F90, Ptotalx.F90, Ptotalxm.F90, Ptotaly.F90, Ptotalym.F90, Ptotalz.F90, and Ptotalzm.F90.

The Green's functions are multiplied according to Eqs 105-107, 111-116, 119-121, 123-128; with the factors inside the integrals.

This function is called from the function kernel.wavenumber.

empymod.kernel.reflections (depth, e\_zH, Gam, lrec, lsrc, use\_ne\_eval)
Calculate Rp, Rm.

 $R_n^{\pm}, \bar{R}_n^{\pm}$ 

This function corresponds to equations 64/65 and A-11/A-12 in [HuTS15], and loosely to the corresponding files Rmin.F90 and Rplus.F90.

This function is called from the function kernel.greenfct.

empymod.kernel.fields (*depth*, *Rp*, *Rm*, *Gam*, *lrec*, *lsrc*, *zsrc*, *ab*, *TM*, *use\_ne\_eval*) Calculate Pu+, Pu-, Pd+, Pd-.

 $P_{s}^{u\pm}, P_{s}^{d\pm}, \bar{P}_{s}^{u\pm}, \bar{P}_{s}^{d\pm}; P_{s-1}^{u\pm}, P_{n}^{u\pm}, \bar{P}_{s-1}^{u\pm}, \bar{P}_{n}^{u\pm}; P_{s+1}^{d\pm}, P_{n}^{d\pm}, \bar{P}_{s+1}^{d\pm}, \bar{P}_{n}^{d\pm}$ 

This function corresponds to equations 81/82, 95/96, 103/104, A-8/A-9, A-24/A-25, and A-32/A-33 in [HuTS15], and loosely to the corresponding files Pdownmin.F90, Pdownplus.F90, Pupmin.F90, and Pdownmin.F90.

This function is called from the function kernel.greenfct.

empymod.kernel.halfspace (off, angle, zsrc, zrec, etaH, etaV, freqtime, ab, signal, solution='dhs') Return frequency- or time-space domain VTI half-space solution.

Calculates the frequency- or time-space domain electromagnetic response for a half-space below air using the diffusive approximation, as given in [SIHM10], where the electric source is located at [0, 0, zsrc], and the electric receiver at [xco, yco, zrec].

It can also be used to calculate the fullspace solution or the separate fields: direct field, reflected field, and airwave; always using the diffusive approximation. See solution-parameter.

This function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and solution parameters.

# 5.5.3 transform - Hankel and Fourier Transforms

Methods to carry out the required Hankel transform from wavenumber to frequency domain and Fourier transform from frequency to time domain.

The functions for the QWE and DLF Hankel and Fourier transforms are based on source files (specified in each function) from the source code distributed with [Key12], which can be found at software.seg.org/2012/0003. These functions are (c) 2012 by Kerry Key and the Society of Exploration Geophysicists, http://software.seg.org/ disclaimer.txt. Please read the NOTICE-file in the root directory for more information regarding the involved licenses.

empymod.transform.fht (zsrc, zrec, lsrc, lrec, off, factAng, depth, ab, etaH, etaV, zetaH, zetaV, xdirect, fhtarg, use ne eval, msrc, mrec)

Hankel Transform using the Digital Linear Filter method.

The *Digital Linear Filter* method was introduced to geophysics by [Ghos70], and made popular and widespread by [Ande75], [Ande79], [Ande82]. The DLF is sometimes referred to as the *Fast Hankel Transform* FHT, from which this routine has its name.

This implementation of the DLF follows [Key12], equation 6. Without going into the mathematical details (which can be found in any of the above papers) and following [Key12], the DLF method rewrites the Hankel transform of the form

$$F(r) = \int_0^\infty f(\lambda) J_v(\lambda r) \, \mathrm{d}\lambda$$

as

$$F(r) = \sum_{i=1}^{n} f(b_i/r)h_i/r ,$$

where h is the digital filter. The Filter abscissae b is given by

$$b_i = \lambda_i r = e^{ai}, \qquad i = -l, -l+1, \cdots, l,$$

with l = (n - 1)/2, and a is the spacing coefficient.

This function is loosely based on get\_CSEM1D\_FD\_FHT.m from the source code distributed with [Key12].

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

Returns

**fEM** [array] Returns frequency-domain EM response. **kcount** [int] Kernel count. For DLF, this is 1. **conv** [bool] Only relevant for QWE/QUAD.

Hankel Transform using Quadrature-With-Extrapolation.

*Quadrature-With-Extrapolation* was introduced to geophysics by [Key12]. It is one of many so-called *ISE* methods to solve Hankel Transforms, where *ISE* stands for Integration, Summation, and Extrapolation.

Following [Key12], but without going into the mathematical details here, the QWE method rewrites the Hankel transform of the form

$$F(r) = \int_0^\infty f(\lambda) J_v(\lambda r) \, \mathrm{d}\lambda$$

as a quadrature sum which form is similar to the DLF (equation 15),

$$F_i \approx \sum_{j=1}^m f(x_j/r) w_j g(x_j) = \sum_{j=1}^m f(x_j/r) \hat{g}(x_j) ,$$

but with various bells and whistles applied (using the so-called Shanks transformation in the form of a routine called  $\epsilon$ -algorithm ([Shan55], [Wynn56]; implemented with algorithms from [Tref00] and [Weni89]).

This function is based on get\_CSEM1D\_FD\_QWE.m, qwe.m, and getBesselWeights.m from the source code distributed with [Key12].

In the spline-version, hqwe checks how steep the decay of the wavenumber-domain result is, and calls QUAD for the very steep interval, for which QWE is not suited.

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

Returns

fEM [array] Returns frequency-domain EM response.kcount [int] Kernel count.conv [bool] If true, QWE/QUAD converged. If not, <htarg> might have to be adjusted.

Hankel Transform using the QUADPACK library.

This routine uses the scipy.integrate.quad module, which in turn makes use of the Fortran library QUADPACK (qagse).

It is massively (orders of magnitudes) slower than either fht or hqwe, and is mainly here for completeness and comparison purposes. It always uses interpolation in the wavenumber domain, hence it generally will not be as precise as the other methods. However, it might work in some areas where the others fail.

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

#### Returns

fEM [array] Returns frequency-domain EM response.kcount [int] Kernel count. For HQUAD, this is 1.conv [bool] If true, QUAD converged. If not, <htarg> might have to be adjusted.

empymod.transform.ffht (fEM, time, freq, ftarg)

Fourier Transform using the Digital Linear Filter method.

It follows the Filter methodology [Ande75], using Cosine- and Sine-filters; see fht for more information.

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

This function is based on get\_CSEM1D\_TD\_FHT.m from the source code distributed with [Key12]. Returns

> **tEM** [array] Returns time-domain EM response of fEM for given time. **conv** [bool] Only relevant for QWE/QUAD.

empymod.transform.fqwe (fEM, time, freq, qweargs)

Fourier Transform using Quadrature-With-Extrapolation.

It follows the QWE methodology [Key12] for the Hankel transform, see hqwe for more information.

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

This function is based on get\_CSEM1D\_TD\_QWE.m from the source code distributed with [Key12].

fqwe checks how steep the decay of the frequency-domain result is, and calls QUAD for the very steep interval, for which QWE is not suited.

# Returns

tEM [array] Returns time-domain EM response of fEM for given time.

 $\textbf{conv} \hspace{0.1 cm} [bool] \hspace{0.1 cm} If \hspace{0.1 cm} true, \hspace{0.1 cm} QWE/QUAD \hspace{0.1 cm} converged. \hspace{0.1 cm} If \hspace{0.1 cm} not, \hspace{0.1 cm} < \hspace{-0.1 cm} tracp \hspace{0.1 cm} might \hspace{0.1 cm} have \hspace{0.1 cm} to \hspace{0.1 cm} be \hspace{0.1 cm} adjusted. \hspace{0.1 cm}$ 

empymod.transform.fftlog(fEM, time, freq, ftarg)

Fourier Transform using FFTLog.

FFTLog is the logarithmic analogue to the Fast Fourier Transform FFT. FFTLog was presented in Appendix B of [Hami00] and published at <a href="http://casa.colorado.edu/~ajsh/FFTLog">http://casa.colorado.edu/~ajsh/FFTLog</a>.

This function uses a simplified version of pyfftlog, which is a python-version of FFTLog. For more details regarding pyfftlog see <a href="https://github.com/prisae/pyfftlog">https://github.com/prisae/pyfftlog</a>.

Not the full flexibility of FFTLog is available here: Only the logarithmic FFT (fftl in FFTLog), not the Hankel transform (fht in FFTLog). Furthermore, the following parameters are fixed:

- kr = 1 (initial value)
- kropt = 1 (silently adjusts kr)
- dir = 1 (forward)

Furthermore, q is restricted to  $-1 \le q \le 1$ .

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

### Returns

**tEM** [array] Returns time-domain EM response of fEM for given time. **conv** [bool] Only relevant for QWE/QUAD.

conv [bool] Only relevant for QwE/QUA

# empymod.transform.fft(fEM, time, freq, ftarg)

Fourier Transform using the Fast Fourier Transform.

The function is called from one of the modelling routines in model. Consult these modelling routines for a description of the input and output parameters.

### Returns

**tEM** [array] Returns time-domain EM response of fEM for given time. **conv** [bool] Only relevant for QWE/QUAD. Digital Linear Filter method.

This is the kernel of the DLF method, used for the Hankel (fht) and the Fourier (ffht) Transforms. See fht for an extensive description.

For the Hankel transform, *signal* contains 3 complex wavenumber-domain signals: (PJ0, PJ1, PJ0b), as returned from *kernel.wavenumber*. The Hankel DLF has two additional, optional parameters: *factAng*, as returned from *kernel.angle\_factor*, and *ab*. The PJ0-kernel is the part of the wavenumber-domain calculation which contains a zeroth-order Bessel function and does NOT depend on the angle between source and receiver, only on offset. PJ0b and PJ1 are the parts of the wavenumber-domain calculation which contain a zeroth- and first-order Bessel function, respectively, and can depend on the angle between source and receiver. PJ0, PJ1, or PJ0b can also be None, if they are not used.

For the Fourier transform, *signal* is a complex frequency-domain signal. The Fourier DLF requires one additional parameter, *kind*, which will be 'cos' or 'sin'.

empymod.transform.qwe (rtol, atol, maxint, inp, intervals, lambd=None, off=None, factAng=None)
Quadrature-With-Extrapolation.

This is the kernel of the QWE method, used for the Hankel (hqwe) and the Fourier (fqwe) Transforms. See hqwe for an extensive description.

This function is based on qwe.m from the source code distributed with [Key12].

empymod.transform.get\_spline\_values (filt, inp, nr\_per\_dec=None)
Return required calculation points.

empymod.transform.**fhti** (*rmin*, *rmax*, *n*, *q*, *mu*) Return parameters required for FFTLog.

# 5.5.4 filters – Digital Linear Filters

Filters for the Digital Linear Filter (DLF) method for the Hankel [Ghos70]) and the Fourier ([Ande75]) transforms.

To calculate the  ${\tt dlf.factor}\ I$  used

np.around(np.average(dlf.base[1:]/dlf.base[:-1]), 15)

The filters kong\_61\_2007 and kong\_241\_2007 from [Kong07], and key\_101\_2009, key\_201\_2009, key\_401\_2009, key\_81\_CosSin\_2009, key\_241\_CosSin\_2009, and key\_601\_CosSin\_2009 from [Key09] are taken from *DIPOLE1D*, [Key09], which can be downloaded at http://marineemlab.ucsd.edu/ Projects/Occam/1DCSEM (1DCSEM). *DIPOLE1D* is distributed under the license GNU GPL version 3 or later. Kerry Key gave his written permission to re-distribute the filters under the Apache License, Version 2.0 (email from Kerry Key to Dieter Werthmüller, 21 November 2016).

The filters and erson\_801\_1982 from [Ande82] and key\_51\_2012, key\_101\_2012, key\_201\_2012, key\_101\_CosSin\_2012, and key\_201\_CosSin\_2012, all from [Key12], are taken from the software distributed with [Key12] and available at http://software.seg.org/2012/0003 (SEG-2012-003). These filters are distributed under the SEG license.

The filter wer\_201\_2018 was designed with the add-on fdesign, see https://github.com/empymod/ article-fdesign.

class empymod.filters.DigitalFilter(name, savename=None)
 Simple Class for Digital Linear Filters.

fromfile (self, path='filters')

Load filter values from ascii-files.

Load filter base and filter coefficients from ascii files in the directory *path*; *path* can be a relative or absolute path.

#### **Examples**

```
>>> import empymod
>>> # Create an empty filter;
>>> # Name has to be the base of the text files
>>> filt = empymod.filters.DigitalFilter('my-filter')
>>> # Load the ascii-files
>>> filt.fromfile()
>>> # This will load the following three files:
>>> # ./filters/my-filter_base.txt
>>> # ./filters/my-filter_j0.txt
>>> # ./filters/my-filter_j1.txt
>>> # and store them in filt.base, filt.j0, and filt.j1.
```

### tofile (self, path='filters')

Save filter values to ascii-files.

Store the filter base and the filter coefficients in separate files in the directory *path*; *path* can be a relative or absolute path.

#### **Examples**

```
>>> import empymod
>>> # Load a filter
>>> filt = empymod.filters.wer_201_2018()
>>> # Save it to pure ascii-files
>>> filt.tofile()
>>> # This will save the following three files:
>>> # ./filters/wer_201_2018_base.txt
>>> # ./filters/wer_201_2018_j0.txt
>>> # ./filters/wer_201_2018_j1.txt
```

## empymod.filters.kong\_61\_2007()

Kong 61 pt Hankel filter, as published in [Kong07].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

#### empymod.filters.kong\_241\_2007()

Kong 241 pt Hankel filter, as published in [Kong07].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

```
empymod.filters.key_101_2009()
```

Key 101 pt Hankel filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

empymod.filters.key\_201\_2009()

Key 201 pt Hankel filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

empymod.filters.key\_401\_2009()
Key 401 pt Hankel filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

```
empymod.filters.anderson_801_1982()
Anderson 801 pt Hankel filter, as published in [Ande82].
```

Taken from file wa801Hankel.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

empymod.filters.key\_51\_2012()
 Key 51 pt Hankel filter, as published in [Key12].

Taken from file kk51Hankel.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

empymod.filters.key\_101\_2012()
Key 101 pt Hankel filter, as published in [Key12].

Taken from file kk101Hankel.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

empymod.filters.key\_201\_2012()
Key 201 pt Hankel filter, as published in [Key12].

Taken from file kk201Hankel.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

empymod.filters.wer\_201\_2018()
Werthmüller 201 pt Hankel filter, 2018.

Designed with the empymod add-on fdesign, see https://github.com/empymod/article-fdesign.

License: Apache License, Version 2.0,.

empymod.filters.key\_81\_CosSin\_2009()
Key 81 pt CosSin filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

empymod.filters.key\_241\_CosSin\_2009()
Key 241 pt CosSin filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

empymod.filters.key\_601\_CosSin\_2009() Key 601 pt CosSin filter, as published in [Key09].

Taken from file FilterModules.f90 provided with 1DCSEM.

License: Apache License, Version 2.0,.

empymod.filters.key\_101\_CosSin\_2012() Key 101 pt CosSin filter, as published in [Key12].

Taken from file kk101CosSin.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

empymod.filters.key\_201\_CosSin\_2012()

Key 201 pt CosSin filter, as published in [Key12].

Taken from file kk201CosSin.txt provided with SEG-2012-003.

License: http://software.seg.org/disclaimer.txt.

# 5.5.5 utils - Utilites

Utilities for model such as checking input parameters.

### This module consists of four groups of functions:

- 0. General settings
- 1. Class EMArray
- 2. Input parameter checks for modelling
- 3. Internal utilities

#### class empymod.utils.EMArray

Subclassing an ndarray: add *amplitude* <amp> and *phase* <pha>.

## Parameters

realpart [array]

- 1. Real part of input, if input is real or complex.
- 2. Imaginary part of input, if input is pure imaginary.
- 3. Complex input.

In cases 2 and 3, imagpart must be None.

imagpart: array, optional Imaginary part of input. Defaults to None.

#### **Examples**

```
>>> import numpy as np
>>> from empymod.utils import EMArray
>>> emvalues = EMArray(np.array([1,2,3]), np.array([1, 0, -1]))
>>> print('Amplitude : ', emvalues.amp)
Amplitude : [ 1.41421356 2. 3.16227766]
>>> print('Phase : ', emvalues.pha)
Phase : [ 45. 0. -18.43494882]
```

#### Attributes

**amp** [ndarray] Amplitude of the input data.

pha [ndarray] Phase of the input data, in degrees, lag-defined (increasing with increasing offset.) To get lead-defined phases, multiply imagpart by -1 before passing through this function.

empymod.utils.check\_time\_only(time, signal, verb)

Check time and signal parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

### Parameters

time [array\_like] Times t (s).
signal [{None, 0, 1, -1}]

### Source signal:

- None: Frequency-domain response
- -1 : Switch-off time-domain response
- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response
- **verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

time [float] Time, checked for size and assured min\_time.

empymod.utils.check\_time(time, signal, ft, ftarg, verb)

Check time domain specific input parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

**time** [array\_like] Times t (s). **signal** [{None, 0, 1, -1}]

#### Source signal:

- None: Frequency-domain response
- -1 : Switch-off time-domain response
- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response

**ft** [{'sin', 'cos', 'qwe', 'fftlog', 'fft'}] Flag for Fourier transform.

ftarg [str or filter from empymod.filters or array\_like,] Only used if signal !=None.

Depends on the value for ft:

**verb** [{0, 1, 2, 3, 4}] Level of verbosity.

## Returns

time [float] Time, checked for size and assured min\_time.

**freq** [float] Frequencies required for given times and ft-settings.

ft, ftarg Checked if valid and set to defaults if not provided, checked with signal.

empymod.utils.check\_model (depth, res, aniso, epermH, epermV, mpermH, mpermV, xdirect,

verb)

Check the model: depth and corresponding layer parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

# Parameters

**depth** [list] Absolute layer interfaces z (m); #depth = #res - 1 (excluding +/- infinity).

**res** [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1.

**aniso** [array\_like] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res.

- **epermH, epermV** [array\_like] Relative horizontal/vertical electric permittivities epsilon\_h/epsilon\_v (-); #epermH = #epermV = #res.
- **mpermH, mpermV** [array\_like] Relative horizontal/vertical magnetic permeabilities mu\_h/mu\_v (-); #mpermH = #mpermV = #res.
- **xdirect** [bool, optional] If True and source and receiver are in the same layer, the direct field is calculated analytically in the frequency domain, if False it is calculated in the wavenumber domain.
- **verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

**depth** [array] Depths of layer interfaces, adds -infty at beginning if not present. **res** [array] As input, checked for size.

aniso [array] As input, checked for size. If None, defaults to an array of ones.

- **epermH, epermV** [array\_like] As input, checked for size. If None, defaults to an array of ones.
- **mpermH, mpermV** [array\_like] As input, checked for size. If None, defaults to an array of ones.
- **isfullspace** [bool] If True, the model is a fullspace (res, aniso, epermH, epermV, mpermM, and mpermV are in all layers the same).

empymod.utils.check\_frequency (freq, res, aniso, epermH, epermV, mpermH, mpermV, verb) Calculate frequency-dependent parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### **Parameters**

freq [array\_like] Frequencies f (Hz).

res [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1.
aniso [array\_like] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res.
epermH, epermV [array\_like] Relative horizontal/vertical electric permittivities epsilon\_h/epsilon\_v (-); #epermH = #epermV = #res.

**mpermH, mpermV** [array\_like] Relative horizontal/vertical magnetic permeabilities mu h/mu v (-); #mpermH = #mpermV = #res.

**verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

freq [float] Frequency, checked for size and assured min\_freq.etaH, etaV [array] Parameters etaH/etaV, same size as provided resistivity.zetaH, zetaV [array] Parameters zetaH/zetaV, same size as provided resistivity.

empymod.utils.check\_hankel(ht, htarg, verb)

Check Hankel transform parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

ht [{'fht', 'qwe', 'quad'}] Flag to choose the Hankel transform.

htarg [str or filter from empymod.filters or array\_like,] Depends on the value for ht.

**verb** [{0, 1, 2, 3, 4}] Level of verbosity.

# Returns

ht, htarg Checked if valid and set to defaults if not provided.

empymod.utils.check\_opt(opt, loop, ht, htarg, verb)

Check optimization parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

## Parameters

 $opt \ [\{None, `parallel'\}] \ Optimization \ flag; use \ \texttt{numexpr} \ or \ not.$ 

**loop** [{None, 'freq', 'off'}] Loop flag.

ht [str] Flag to choose the Hankel transform.

htarg [array\_like,] Depends on the value for ht.

**verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

use\_ne\_eval [bool] Boolean if to use numexpr. loop\_freq [bool] Boolean if to loop over frequencies. loop\_off [bool] Boolean if to loop over offsets.

empymod.utils.check\_dipole(inp, name, verb)

Check dipole parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

inp [list of floats or arrays] Pole coordinates (m): [pole-x, pole-y, pole-z].
name [str, {'src', 'rec'}] Pole-type.
verb [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

**inp** [list] List of pole coordinates [x, y, z]. **ninp** [int] Number of inp-elements

### empymod.utils.check\_bipole(inp, name)

Check di-/bipole parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

inp [list of floats or arrays] Coordinates of inp (m): [dipole-x, dipole-y, dipole-z, az-imuth, dip] or. [bipole-x0, bipole-x1, bipole-y0, bipole-y1, bipole-z0, bipole-z1].
 name [str, {'src', 'rec'}] Pole-type.

#### Returns

inp [list] As input, checked for type and length.ninp [int] Number of inp.ninpz [int] Number of inp depths (ninpz is either 1 or ninp).isdipole [bool] True if inp is a dipole.

# empymod.utils.check\_ab(ab, verb)

Check source-receiver configuration.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

**ab** [int] Source-receiver configuration. **verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

**ab\_calc** [int] Adjusted source-receiver configuration using reciprocity. **msrc, mrec** [bool] If True, src/rec is magnetic; if False, src/rec is electric.

empymod.utils.check\_solution(solution, signal, ab, msrc, mrec)

Check required solution with parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

# Parameters

**solution** [str] String to define analytical solution. **signal** [{None, 0, 1, -1}]

#### Source signal:

- None: Frequency-domain response
- -1 : Switch-off time-domain response
- 0 : Impulse time-domain response
- +1 : Switch-on time-domain response

msrc, mrec [bool] True if src/rec is magnetic, else False.

empymod.utils.get\_abs (msrc, mrec, srcazm, srcdip, recazm, recdip, verb)
Get required ab's for given angles.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

**msrc, mrec** [bool] True if src/rec is magnetic, else False. **srcazm, recazm** [float] Horizontal source/receiver angle (azimuth). **srcdip, recdip** [float] Vertical source/receiver angle (dip). **verb** [{0, 1, 2, 3, 4}] Level of verbosity.

## Returns

**ab\_calc** [array of int] ab's to calculate for this bipole.

empymod.utils.get\_geo\_fact (ab, srcazm, srcdip, recazm, recdip, msrc, mrec)
Get required geometrical scaling factor for given angles.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

ab [int] Source-receiver configuration.srcazm, recazm [float] Horizontal source/receiver angle.srcdip, recdip [float] Vertical source/receiver angle.

### Returns

fact [float] Geometrical scaling factor.

empymod.utils.get\_azm\_dip (*inp*, *iz*, *ninpz*, *intpts*, *isdipole*, *strength*, *name*, *verb*) Get angles, interpolation weights and normalization weights. This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

### **Parameters**

inp [list of floats or arrays]

#### Input coordinates (m):

- [x0, x1, y0, y1, z0, z1] (bipole of finite length)
- [x, y, z, azimuth, dip] (dipole, infinitesimal small)

iz [int] Index of current di-/bipole depth (-).
ninpz [int] Total number of di-/bipole depths (ninpz = 1 or npinz = nsrc) (-).
intpts [int] Number of integration points for bipole (-).
isdipole [bool] Boolean if inp is a dipole.
strength [float, optional]

#### Source strength (A):

- If 0, output is normalized to source and receiver of 1 m length, and source strength of 1 A.
- If != 0, output is returned for given source and receiver length, and source strength.

**name** [str, {'src', 'rec'}] Pole-type.

**verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

tout [list of floats or arrays] Dipole coordinates x, y, and z (m).

azm [float or array of floats] Horizontal angle (azimuth).

dip [float or array of floats] Vertical angle (dip).

g\_w [float or array of floats] Factors from Gaussian interpolation.

intpts [int] As input, checked.

inp\_w [float or array of floats] Factors from source/receiver length and source strength.

### empymod.utils.get\_off\_ang(src, rec, nsrc, nrec, verb)

Get depths, offsets, angles, hence spatial input parameters.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

**src, rec** [list of floats or arrays] Source/receiver dipole coordinates x, y, and z (m). **nsrc, nrec** [int] Number of sources/receivers (-). **verb** [{0, 1, 2, 3, 4}] Level of verbosity.

#### Returns

**off** [array of floats] Offsets **angle** [array of floats] Angles

empymod.utils.get\_layer\_nr(inp, depth)

Get number of layer in which inp resides.

Note: If zinp is on a layer interface, the layer above the interface is chosen.

This check-function is called from one of the modelling routines in model. Consult these modelling routines for a detailed description of the input parameters.

#### Parameters

**inp** [list of floats or arrays] Dipole coordinates (m) **depth** [array] Depths of layer interfaces.

#### Returns

**linp** [int or array\_like of int] Layer number(s) in which inp resides (plural only if bipole). **zinp** [float or array] inp[2] (depths).

empymod.utils.printstartfinish (verb, inp=None, kcount=None)
Print start and finish with time measure and kernel count.

empymod.utils.conv\_warning (conv, targ, name, verb) Print error if QWE/QUAD did not converge at least once. Set minimum values of parameters.

The given parameters are set to its minimum value if they are smaller.

# Parameters

- min\_freq [float, optional] Minimum frequency [Hz] (default 1e-20 Hz).
- min\_time [float, optional] Minimum time [s] (default 1e-20 s).
- **min\_off** [float, optional] Minimum offset [m] (default 1e-3 m). Also used to round src-& rec-coordinates.
- **min\_res** [float, optional] Minimum horizontal and vertical resistivity [Ohm.m] (default 1e-20).

min\_angle [float, optional] Minimum angle [-] (default 1e-10).

#### empymod.utils.get\_minimum()

Return the current minimum values.

### Returns

min\_vals [dict] Dictionary of current minimum values with keys

- min\_freq : float
- min\_time : float
- min\_off : float
- min\_res : float
- min\_angle : float

For a full description of these options, see set\_minimum.

empymod.utils.spline\_backwards\_hankel(ht, htarg, opt)

Check opt if deprecated 'spline' is used.

Returns corrected htarg, opt. r

```
class empymod.utils.Report (add_pckg=None, ncol=3, text_width=80, sort=False)
Print date, time, and version information.
```

Use scooby to print date, time, and package version information in any environment (Jupyter notebook, IPython console, Python console, QT console), either as html-table (notebook) or as plain text (anywhere).

Always shown are the OS, number of CPU(s), numpy, scipy, empymod, sys.version, and time/date.

Additionally shown are, if they can be imported, numexpr, IPython, and matplotlib. It also shows MKL information, if available.

All modules provided in add\_pckg are also shown.

#### Parameters

- **add\_pckg** [packages, optional] Package or list of packages to add to output information (must be imported beforehand).
- **ncol** [int, optional] Number of package-columns in html table (no effect in text-version); Defaults to 3.

text\_width [int, optional] The text width for non-HTML display modes

sort [bool, optional] Sort the packages when the report is shown

# Examples

**class** empymod.utils.**Versions** (*add\_pckg=None*, *ncol=3*) New name is *Report*, here for backwards compatibility.

empymod.utils.versions(mode=None, add\_pckg=None, ncol=4)
Old func-way of class Report, here for backwards compatibility.

# 5.6 Add-ons

# 5.6.1 fdesign – Digital Linear Filter (DLF) design

The add-on fdesign can be used to design digital linear filters for the Hankel or Fourier transform, or for any linear transform ([Ghos70]). For this included or provided theoretical transform pairs can be used. Alternatively, one can use the EM modeller empymod to use the responses to an arbitrary 1D model as numerical transform pair.

More information can be found in the following places:

- The article about fdesign is in the repo https://github.com/empymod/article-fdesign
- Example notebooks to design a filter can be found in the repo https://github.com/empymod/ empymod-examples

This filter designing tool uses the direct matrix inversion method as described in [Kong07] and is based on scripts by [Key12]. The whole project of fdesign started with the Matlab scripts from Kerry Key, which he used to design his filters for [Key09], [Key12]. Fruitful discussions with Evert Slob and Kerry Key improved the add-on substantially.

Note that the use of empymod to create numerical transform pairs is, as of now, only implemented for the Hankel transform.

### Implemented analytical transform pairs

The following tables list the transform pairs which are implemented by default. Any other transform pair can be provided as input. A transform pair is defined in the following way:

```
from empymod.scripts import fdesign

def my_tp_pair(var):
    '''My transform pair.'''

    def lhs(l):
        return func(l, var)

    def rhs(r):
        return func(r, var)

    return fdesign.Ghosh(name, lhs, rhs)
```

Here, name must be one of j0, j1, sin, or cos, depending what type of transform pair it is. Additional variables are provided with var. The evaluation points of the lhs are denoted by 1, and the evaluation points of the rhs are denoted as r. As an example here the implemented transform pair  $j0_1$ 

```
def j0_1(a=1):
    '''Hankel transform pair J0_1 ([Ande75]_).'''
    def lhs(l):
        return l*np.exp(-a*l**2)
    def rhs(r):
        return np.exp(-r**2/(4*a))/(2*a)
```

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return Ghosh('j0', lhs, rhs)

# Implemented Hankel transforms

• j0\_1 [Ande75]

$$\int_0^\infty l \exp\left(-al^2\right) J_0(lr) dl = \frac{\exp\left(\frac{-r^2}{4a}\right)}{2a}$$

• j0\_2 [Ande75]

$$\int_{0}^{\infty} \exp(-al) J_{0}(lr) dl = \frac{1}{\sqrt{a^{2} + r^{2}}}$$

• j0\_3 [GuSi97]

$$\int_0^\infty l \exp(-al) J_0(lr) dl = \frac{a}{(a^2 + r^2)^{3/2}}$$

• j0\_4 [ChCo82]

$$\int_0^\infty \frac{l}{\beta} \exp\left(-\beta z_v\right) J_0(lr) dl = \frac{\exp\left(-\gamma R\right)}{R}$$

• j0\_5 [ChCo82]

$$\int_0^\infty l \exp\left(-\beta z_v\right) J_0(lr) dl = \frac{z_v(\gamma R+1)}{R^3} \exp\left(-\gamma R\right)$$

• j1\_1 [Ande75]

$$\int_0^\infty l^2 \exp\left(-al^2\right) J_1(lr) dl = \frac{r}{4a^2} \exp\left(-\frac{r^2}{4a}\right)$$

• j1\_2 [Ande75]

$$\int_0^\infty \exp(-al) J_1(lr) dl = \frac{\sqrt{a^2 + r^2} - a}{r\sqrt{a^2 + r^2}}$$

• j1\_3 [Ande75]

$$\int_0^\infty l \exp(-al) J_1(lr) dl = \frac{r}{(a^2 + r^2)^{3/2}}$$

• j1\_4 [ChCo82]

$$\int_0^\infty \frac{l^2}{\beta} \exp\left(-\beta z_v\right) J_1(lr) dl = \frac{r(\gamma R+1)}{R^3} \exp\left(-\gamma R\right)$$

• j1\_5 [ChCo82]

$$\int_{0}^{\infty} l^{2} \exp(-\beta z_{v}) J_{1}(lr) dl = \frac{r z_{v}(\gamma^{2} R^{2} + 3\gamma R + 3)}{R^{5}} \exp(-\gamma R)$$

Where

$$a > 0, r > 0$$
$$z_v = |z_{rec} - z_{src}|$$
$$R = \sqrt{r^2 + z_v^2}$$
$$\gamma = \sqrt{2j\pi\mu_0 f/\rho}$$
$$\beta = \sqrt{l^2 + \gamma^2}$$

### **Implemented Fourier transforms**

• sin\_1 [Ande75]

$$\int_0^\infty l \exp\left(-a^2 l^2\right) \sin(lr) dl = \frac{\sqrt{\pi}r}{4a^3} \exp\left(-\frac{r^2}{4a^2}\right)$$

• sin\_2 [Ande75]

$$\int_0^\infty \exp\left(-al\right) \sin(lr) dl = \frac{r}{a^2 + r^2}$$

• sin\_3 [Ande75]

$$\int_0^\infty \frac{l}{a^2 + l^2} \sin(lr) dl = \frac{\pi}{2} \exp\left(-ar\right)$$

• cos\_1 [Ande75]

$$\int_0^\infty \exp\left(-a^2l^2\right)\cos(lr)dl = \frac{\sqrt{\pi}}{2a}\exp\left(-\frac{r^2}{4a^2}\right)$$

• cos\_2 [Ande75]

$$\int_0^\infty \exp\left(-al\right)\cos(lr)dl = \frac{a}{a^2 + r^2}$$

• cos\_3 [Ande75]

$$\int_0^\infty \frac{1}{a^2 + l^2} \cos(lr) dl = \frac{\pi}{2a} \exp\left(-ar\right)$$

empymod.scripts.fdesign.design (n, spacing, shift, fl, fC=False, r=None, r\_def=(1, 1, 2), reim=None, cvar='amp', error=0.01, name=None, full\_output=False, finish=False, save=True, path='filters', verb=2, plot=1)

Digital linear filter (DLF) design

This routine can be used to design digital linear filters for the Hankel or Fourier transform, or for any linear transform ([Ghos70]). For this included or provided theoretical transform pairs can be used. Alternatively, one can use the EM modeller empymod to use the responses to an arbitrary 1D model as numerical transform pair.

This filter designing tool uses the direct matrix inversion method as described in [Kong07] and is based on scripts by [Key12]. The tool is an add-on to the electromagnetic modeller empymod [Wert17]. Fruitful discussions with Evert Slob and Kerry Key improved the add-on substantially.

Example notebooks of its usage can be found in the repo github.com/empymod/empymod-examples. **Parameters** 

- **n** [int] Filter length.
- **spacing: float or tuple (start, stop, num)** Spacing between filter points. If tuple, it corresponds to the input for np.linspace with endpoint=True.
- **shift: float or tuple (start, stop, num)** Shift of base from zero. If tuple, it corresponds to the input for np.linspace with endpoint=True.
- **fI**, **fC** [transform pairs] Theoretical or numerical transform pair(s) for the inversion (I) and for the check of goodness (fC). fC is optional. If not provided, fI is used for both fI and fC.
- **r** [array, optional] Right-hand side evaluation points for the check of goodness (fC). Defaults to r = np.logspace(0, 5, 1000), which are a lot of evaluation points, and depending on the transform pair way too long r's.
- **r\_def** [tuple (add\_left, add\_right, factor), optional] Definition of the right-hand side evaluation points r of the inversion. r is derived from the base values, default is (1, 1, 2).
  - $rmin = log10(1/max(base)) add_left$
  - rmax = log10(1/min(base)) + add\_right
  - r = logspace(rmin, rmax, factor\*n)
- **reim** [np.real or np.imag, optional] Which part of complex transform pairs is used for the inversion. Defaults to np.real.
- **cvar** [string {'amp', 'r'}, optional] If 'amp', the inversion minimizes the amplitude. If 'r', the inversion maximizes the right-hand side evaluation point r. Defaults is 'amp'.
- **error** [float, optional] Up to which relative error the transformation is considered good in the evaluation of the goodness. Default is 0.01 (1 %).
- name [str, optional] Name of the filter. Defaults to dlf\_+str(n).
- **full\_output** [bool, optional] If True, returns best filter and output from scipy.optimize.brute; else only filter. Default is False.
- **finish** [None, True, or callable, optional] If callable, it is passed through to scipy.optimize.brute: minimization function to find minimize best result from brute-force approach. Default is None. You can simply provide True in order to use scipy.optimize.fmin\_powell(). Set this to None if you are only interested in the actually provided spacing/shift-values.
- save [bool, optional] If True, best filter is saved to plain text files in ./filters/. Can be loaded with fdesign.load\_filter(name). If full, the inversion output is stored too. You can add '.gz' to *name*, which will then save the full inversion output in a compressed file instead of plain text.
- **path** [string, optional] Absolute or relative path where output will be saved if *save=True*. Default is 'filters'.
- **verb** [{0, 1, 2}, optional]

#### Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.
- 2: Print additional time, progress, and result
- **plot** [{0, 1, 2, 3}, optional]

#### Level of plot-verbosity, default is 1:

- 0: Plot nothing.
- 1: Plot brute-force result
- 2: Plot additional theoretical transform pairs, and best inv.
- 3: Plot additional inversion result (can result in lots of plots depending on spacing and shift) If you are using a notebook, use %matplotlib notebook to have all inversion results appear in the same plot.

Returns

filter [empymod.filter.DigitalFilter instance] Best filter for the input parameters.
full [tuple] Output from scipy.optimize.brute with full\_output=True. (Returned when
full\_output is True.)

empymod.scripts.fdesign.save\_filter(name, filt, full=None, path='filters')
Save DLF-filter and inversion output to plain text files.

empymod.scripts.fdesign.load\_filter (*name*, *full=False*, *path='filters'*) Load saved DLF-filter and inversion output from text files.

empymod.scripts.fdesign.plot\_result (filt, full, prntres=True)
 QC the inversion result.
 Parameters

- filt, full as returned from fdesign.design with full\_output=True

- If prntres is True, it calls fdesign.print\_result as well.

r

empymod.scripts.fdesign.print\_result (filt, full=None)
Print best filter information.

Parameters

### - filt, full as returned from fdesign.design with full\_output=True

**class** empymod.scripts.fdesign.**Ghosh** (*name*, *lhs*, *rhs*) Simple Class for Theoretical Transform Pairs.

Named after D. P. Ghosh, honouring his 1970 Ph.D. thesis with which he introduced the digital filter method to geophysics ([Ghos70]).

empymod.scripts.fdesign.j0\_1(a=1)
Hankel transform pair J0\_1 ([Ande75]).

empymod.scripts.fdesign.j0\_2(a=1)
Hankel transform pair J0\_2 ([Ande75]).

```
empymod.scripts.fdesign.j0_3(a=1)
Hankel transform pair J0_3 ([GuSi97]).
```

```
empymod.scripts.fdesign.j0_4 (f=1, rho=0.3, z=50)
Hankel transform pair J0_4 ([ChCo82]).
```

#### Parameters

f [float] Frequency (Hz)
rho [float] Resistivity (Ohm.m)
z [float] Vertical distance between source and receiver (m)

```
empymod.scripts.fdesign.j0_5(f=1, rho=0.3, z=50)
```

```
Hankel transform pair J0_5 ([ChCo82]).
```

# Parameters

f [float] Frequency (Hz)
rho [float] Resistivity (Ohm.m)
z [float] Vertical distance between source and receiver (m)

empymod.scripts.fdesign.j1\_1(a=1)
Hankel transform pair J1\_1 ([Ande75]).

empymod.scripts.fdesign.j1\_2(*a=1*) Hankel transform pair J1\_2([Ande75]).

empymod.scripts.fdesign.j1\_3(a=1)
Hankel transform pair J1\_3 ([Ande75]).

empymod.scripts.fdesign.j1\_4 (*f*=1, *rho*=0.3, *z*=50) Hankel transform pair J1\_4 ([ChCo82]). Parameters f [float] Frequency (Hz)

**rho** [float] Resistivity (Ohm.m)

```
z [float] Vertical distance between source and receiver (m)
empymod.scripts.fdesign.j1_5(f=1, rho=0.3, z=50)
     Hankel transform pair J1_5 ([ChCo82]).
          Parameters
                f [float] Frequency (Hz)
                rho [float] Resistivity (Ohm.m)
               z [float] Vertical distance between source and receiver (m)
empymod.scripts.fdesign.sin_1(a=1)
     Fourier sine transform pair sin_1 ([Ande75]).
empymod.scripts.fdesign.sin 2(a=1)
     Fourier sine transform pair sin 2 ([Ande75]).
empymod.scripts.fdesign.sin_3(a=1)
     Fourier sine transform pair sin_3 ([Ande75]).
empymod.scripts.fdesign.cos_1(a=1)
     Fourier cosine transform pair cos_1 ([Ande75]).
empymod.scripts.fdesign.cos_2(a=1)
     Fourier cosine transform pair cos_2 ([Ande75]).
empymod.scripts.fdesign.cos_3(a=1)
     Fourier cosine transform pair cos 3 ([Ande75]).
empymod.scripts.fdesign.empy_hankel (ftype, zsrc, zrec, res, freqtime, depth=None,
                                                             epermH=None,
                                                                              epermV=None,
                                              aniso=None,
                                              mpermH=None,
                                                               mpermV=None,
                                                                                htarg=None,
                                              verblhs=0, verbrhs=0)
```

Numerical transform pair with empymod.

All parameters except ftype, verblhs, and verbrhs correspond to the input parameters to empymod. dipole. See there for more information.

Note that if depth=None or [], the analytical full-space solutions will be used (much faster).

Parameters

ftype [str or list of strings] Either of: {'j0', 'j1', 'j2', ['j0', 'j1']}

- 'j0': Analyze J0-term with ab=11, angle=45°
- 'j1': Analyze J1-term with ab=31, angle=0°
- 'j2': Analyze J0- and J1-terms jointly with ab=12, angle=45°

• ['j0', 'j1']: Same as calling empy\_hankel twice, once with 'j0' and one with 'j1'; can be provided like this to fdesign.design. verblhs, verbrhs: int verb-values provided to empymod for lhs and rhs.

Note that ftype='j2' only works for fC, not for fI.

# 5.6.2 tmtemod - Calculate up- and down-going TM and TE modes

This add-on for empymod adjusts [HuTS15] for TM/TE-split. The development was initiated by the development of https://github.com/empymod/csem-ziolkowski-and-slob ([ZiS119]).

This is a stripped-down version of empymod with a lot of simplifications but an important addition. The modeller empymod returns the total field, hence not distinguishing between TM and TE mode, and even less between up and down-going fields. The reason behind this is simple: The derivation of [HuTS15], on which empymod is based, returns the total field. In this derivation each mode (TM and TE) contains non-physical contributions. The non-physical contributions have opposite signs in TM and TE, so they cancel each other out in the total field. However, in order to obtain the correct TM and TE contributions one has to remove these non-physical parts.

This is what this routine does, but only for an x-directed electric source with an x-directed electric receiver, and in the frequency domain (src and rec in same layer). This version of dipole returns the signal separated into

TM++, TM+-, TM-+, TM-, TE++, TE+-, TE++, and TE- as well as the direct field TM and TE contributions. The first superscript denotes the direction in which the field diffuses towards the receiver and the second superscript denotes the direction in which the field diffuses away from the source. For both the plus-sign indicates the field diffuses in the downward direction and the minus-sign indicates the field diffuses in the upward direction. It uses empymod wherever possible. See the corresponding functions in empymod for more explanation and documentation regarding input parameters. There are important limitations:

- ab == 11 [=> x-directed el. source & el. receivers]
- signal == None [=> only frequency domain]
- xdirect == False [=> direct field calc. in wavenr-domain]
- ht == 'fht'
- htarg == 'key\_201\_2012'
- Options ft, ftarg, opt, and loop are not available.
- lsrc == lrec [=> src & rec are assumed in same layer!]
- Model must have more than 1 layer
- Electric permittivity and magnetic permeability are isotropic.
- Only one frequency at once.

# Theory

The derivation of [HuTS15], on which empymod is based, returns the total field. Internally it also calculates TM and TE modes, and sums these up. However, the separation into TM and TE mode introduces a singularity at  $\kappa = 0$ . It has no contribution in the space-frequency domain to the total fields, but it introduces non-physical events in each mode with opposite sign (so they cancel each other out in the total field). In order to obtain the correct TM and TE contributions one has to remove these non-physical parts.

To remove the non-physical part we use the file tmtemod.py in this directory. This routine is basically a heavily simplified version of empymod with the following limitations outlined above.

So tmtemod.py returns the signal separated into TM++, TM+-, TM-+, TM-, TE++, TE+-, TE-+, and TE- as well as the direct field TM and TE contributions. The first superscript denotes the direction in which the field diffuses towards the receiver and the second superscript denotes the direction in which the field diffuses away from the source. For both the plus-sign indicates the field diffuses in the downward direction and the minus-sign indicates the field diffuses in the upward direction. The routine uses empymod wherever possible, see the corresponding functions in empymod for more explanation and documentation regarding input parameters.

Please note that the notation in [HuTS15] differs from the notation in [ZiS119]. I specify therefore always, which notification applies, either *Hun15* or *Zio19*.

We start with equation (105) in *Hun15*:

$$\hat{G}_{xx}^{ee}(\boldsymbol{x}, \boldsymbol{x}', \omega) = \hat{G}_{xx;s}^{ee;i}(\boldsymbol{x} - \boldsymbol{x}', \omega) + \frac{1}{8\pi} \int_{\kappa=0}^{\infty} \left( \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} - \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} \right) J_0(\kappa r) \kappa d\kappa$$
$$- \frac{\cos(2\phi)}{8\pi} \int_{\kappa=0}^{\infty} \left( \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} + \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} \right) J_2(\kappa r) \kappa d\kappa.$$

Ignoring the incident field, and using  $J_2 = \frac{2}{\kappa r} J_1 - J_0$  to avoid  $J_2$ -integrals, we get

$$\hat{G}_{xx}^{ee}(\boldsymbol{x}, \boldsymbol{x}', \omega) = \frac{1}{8\pi} \int_{\kappa=0}^{\infty} \left( \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} - \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} \right) J_0(\kappa r) \kappa \,\mathrm{d}\kappa$$
$$+ \frac{\cos(2\phi)}{8\pi} \int_{\kappa=0}^{\infty} \left( \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} + \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} \right) J_0(\kappa r) \kappa \,\mathrm{d}\kappa$$

$$-\frac{\cos(2\phi)}{4\pi r}\int_{\kappa=0}^{\infty}\left(\frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} + \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s}\right) J_1(\kappa r) \,\mathrm{d}\kappa.$$

From this the TM- and TE-parts follow as

$$TE = \frac{\cos(2\phi) - 1}{8\pi} \int_{\kappa=0}^{\infty} \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} J_0(\kappa r) \kappa \, \mathrm{d}\kappa - \frac{\cos(2\phi)}{4\pi r} \int_{\kappa=0}^{\infty} \frac{\zeta_s \tilde{g}_{zz;s}^{te}}{\bar{\Gamma}_s} J_1(\kappa r) \, \mathrm{d}\kappa,$$
  
$$TM = \frac{\cos(2\phi) + 1}{8\pi} \int_{\kappa=0}^{\infty} \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} J_0(\kappa r) \kappa \, \mathrm{d}\kappa - \frac{\cos(2\phi)}{4\pi r} \int_{\kappa=0}^{\infty} \frac{\Gamma_s \tilde{g}_{hh;s}^{tm}}{\eta_s} J_1(\kappa r) \, \mathrm{d}\kappa.$$

Equations (108) and (109) in Hun15 yield the required parameters  $\tilde{g}_{hh;s}^{tm}$  and  $\tilde{g}_{zz;s}^{te}$ ,

$$\begin{split} \tilde{g}_{hh;s}^{tm} &= P_s^{u-} W_s^u + P_s^{d-} W_s^d, \\ \tilde{g}_{zz;s}^{te} &= \bar{P}_s^{u+} \bar{W}_s^u + \bar{P}_s^{d+} \bar{W}_s^d \; . \end{split}$$

The parameters  $P_s^{u\pm}$  and  $P_s^{d\pm}$  are given in equations (81) and (82),  $\bar{P}_s^{u\pm}$  and  $\bar{P}_s^{d\pm}$  in equations (A-8) and (A-9);  $W_s^u$  and  $W_s^d$  in equation (74) in Hun15. This yields

$$\begin{split} \tilde{g}_{zz;s}^{te} &= \frac{R_s^+}{\bar{M}_s} \left\{ \exp[-\bar{\Gamma}_s(z_s - z + d^+)] + \bar{R}_s^- \exp[-\bar{\Gamma}_s(z_s - z + d_s + d^-)] \right\} \\ &+ \frac{\bar{R}_s^-}{\bar{M}_s} \left\{ \exp[-\bar{\Gamma}_s(z - z_{s-1} + d^-)] + \bar{R}_s^+ \exp[-\bar{\Gamma}_s(z - z_{s-1} + d_s + d^+)] \right\}, \\ &= \frac{\bar{R}_s^+}{\bar{M}_s} \left\{ \exp[-\bar{\Gamma}_s(2z_s - z - z')] + \bar{R}_s^- \exp[-\bar{\Gamma}_s(z' - z + 2d_s)] \right\} \\ &+ \frac{\bar{R}_s^-}{\bar{M}_s} \left\{ \exp[-\bar{\Gamma}_s(z + z' - 2z_{s-1})] + \bar{R}_s^+ \exp[-\bar{\Gamma}_s(z - z' + 2d_s)] \right\}, \end{split}$$

where  $d^{\pm}$  is taken from the text below equation (67). There are four terms in the right-hand side, two in the first line and two in the second line. The first term in the first line is the integrand of TE+-, the second term in the first line corresponds to TE++, the first term in the second line is TE-+, and the second term in the second line is TE-.

If we look at TE+-, we have

$$\tilde{g}_{zz;s}^{te+-} = \frac{\bar{R}_s^+}{\bar{M}_s} \exp[-\bar{\Gamma}_s(2z_s - z - z')],$$

and therefore

$$TE^{+-} = \frac{\cos(2\phi) - 1}{8\pi} \int_{\kappa=0}^{\infty} \frac{\zeta_s R_s^+}{\bar{\Gamma}_s \bar{M}_s} \exp[-\bar{\Gamma}_s (2z_s - z - z')] J_0(\kappa r) \kappa \, \mathrm{d}\kappa$$
$$-\frac{\cos(2\phi)}{4\pi r} \int_{\kappa=0}^{\infty} \frac{\zeta_s \bar{R}_s^+}{\bar{\Gamma}_s \bar{M}_s} \exp[-\bar{\Gamma}_s (2z_s - z - z')] J_1(\kappa r) \, \mathrm{d}\kappa.$$

We can compare this to equation (4.165) in Zio19, with  $\hat{I}_x^e = 1$  and slightly re-arranging it to look more alike, we get

$$\hat{E}_{xx;H}^{+-} = \frac{y^2}{4\pi r^2} \int_{\kappa=0}^{\infty} \frac{\zeta_1}{\Gamma_1} \frac{R_{H;1}^-}{M_{H;1}} \exp(-\Gamma_1 h^{+-}) J_0(\kappa r) \kappa d\kappa$$
$$+ \frac{x^2 - y^2}{4\pi r^3} \int_{\kappa=0}^{\infty} \frac{\zeta_1}{\Gamma_1} \left( \frac{R_{H;1}^-}{M_{H;1}} - \frac{R_{H;1}^-(\kappa=0)}{M_{H;1}(\kappa=0)} \right) \exp(-\Gamma_1 h^{+-}) J_1(\kappa r) d\kappa$$
$$- \frac{\zeta_1 (x^2 - y^2)}{4\pi \gamma_1 r^4} \frac{R_{H;1}^-(\kappa=0)}{M_{H;1}(\kappa=0)} \exp(-\gamma_1 R^{+-}).$$

The notation in this equation follows Zio19.

The difference between the two previous equations is that the first one contains non-physical contributions. These have opposite signs in TM+- and TE+-, and therefore cancel each other out. But if we want to know the specific

contributions from TM and TE we have to remove them. The non-physical contributions only affect the  $J_1$ -integrals, and only for  $\kappa = 0$ .

The following lists for all 8 cases the term that has to be removed, in the notation of Zio19 (for the notation as in Hun15 see the implementation in tmtemod.py):

$$\begin{split} TE^{++} &= + \frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{\exp(-\gamma_1|h^-|)}{M_{H;1}(\kappa = 0)},\\ TE^{-+} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{H;1}^+(\kappa = 0)\exp(-\gamma_1 h^{-+})}{M_{H;1}(\kappa = 0)},\\ TE^{+-} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{H;1}^-(\kappa = 0)\exp(-\gamma_1 h^{+-})}{M_{H;1}(\kappa = 0)},\\ TE^{--} &= +\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{H;1}^+(\kappa = 0)R_{H;1}^-(\kappa = 0)\exp(-\gamma_1 h^{--})}{M_{H;1}(\kappa = 0)},\\ TM^{++} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{V;1}^+(\kappa = 0)\exp(-\gamma_1 h^{-+})}{M_{V;1}(\kappa = 0)},\\ TM^{-+} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{V;1}^+(\kappa = 0)\exp(-\gamma_1 h^{-+})}{M_{V;1}(\kappa = 0)},\\ TM^{+-} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{V;1}^-(\kappa = 0)\exp(-\gamma_1 h^{+-})}{M_{V;1}(\kappa = 0)},\\ TM^{--} &= -\frac{\zeta_1(x^2 - y^2)}{4\pi\gamma_1 r^4} \frac{R_{V;1}^+(\kappa = 0)R_{V;1}^-(\kappa = 0)\exp(-\gamma_1 h^{--})}{M_{V;1}(\kappa = 0)}. \end{split}$$

Note that in the first and fourth equations the correction terms have opposite sign as those in the fifth and eighth equations because at  $\kappa = 0$  the TM and TE mode correction terms are equal. Also note that in the second and third equations the correction terms have the same sign as those in the sixth and seventh equations because at  $\kappa = 0$  the TM and TE mode reflection responses in those terms are equal but with opposite sign:  $R_{V;1}^{\pm}(\kappa = 0) = -R_{V;1}^{\pm}(\kappa = 0)$ .

Hun15 uses  $\phi$ , whereas Zio19 uses x, y, for which we can use

$$\cos(2\phi) = -\frac{x^2 - y^2}{r^2}.$$

Return the electromagnetic field due to a dipole source.

This is a modified version of empymod.model.dipole(). It returns the separated contributions of TM-, TM-+, TM+-, TM++, TMdirect, TE-, TE++, TE++, and TEdirect.

#### Parameters

**src, rec** [list of floats or arrays] Source and receiver coordinates (m): [x, y, z]. The xand y-coordinates can be arrays, z is a single value. The x- and y-coordinates must have the same dimension.

Sources or receivers placed on a layer interface are considered in the upper layer.

Sources and receivers must be in the same layer.

- **depth** [list] Absolute layer interfaces z (m); #depth = #res 1 (excluding +/- infinity). **res** [array\_like] Horizontal resistivities rho\_h (Ohm.m); #res = #depth + 1.
- **freqtime** [float] Frequency f (Hz). (The name freqtime is kept for consistency with

empymod.model.dipole(). Only one frequency at once.

- **aniso** [array\_like, optional] Anisotropies lambda = sqrt(rho\_v/rho\_h) (-); #aniso = #res. Defaults to ones.
- **eperm** [array\_like, optional] Relative electric permittivities epsilon (-); #eperm = #res. Default is ones.

- **mperm** [array\_like, optional] Relative magnetic permeabilities mu (-); #mperm = #res. Default is ones.
- **verb** [{0, 1, 2, 3, 4}, optional]

## Level of verbosity, default is 2:

- 0: Print nothing.
- 1: Print warnings.
- 2: Print additional runtime and kernel calls
- 3: Print additional start/stop, condensed parameter information.
- 4: Print additional full parameter information

# Returns

**TM**, **TE** [list of ndarrays, (nfreq, nrec, nsrc)] Frequency-domain EM field [V/m], separated into TM = [TM–, TM++, TM++, TMdirect] and TE = [TE–, TE++, TE++, TE++, TEdirect].

However, source and receiver are normalised. So the source strength is 1 A and its length is 1 m. Therefore the electric field could also be written as [V/(A.m2)].

The shape of EM is (nfreq, nrec, nsrc). However, single dimensions are removed.

# 5.7 References

# Bibliography

- [Ande75] Anderson, W. L., 1975, Improved digital filters for evaluating Fourier and Hankel transform integrals: USGS, PB242800; pubs.er.usgs.gov/publication/70045426.
- [Ande79] Anderson, W. L., 1979, Numerical integration of related Hankel transforms of orders 0 and 1 by adaptive digital filtering: Geophysics, 44, 1287–1305; DOI: 10.1190/1.1441007.
- [Ande82] Anderson, W. L., 1982, Fast Hankel transforms using related and lagged convolutions: ACM Trans. on Math. Softw. (TOMS), 8, 344–368; DOI: 10.1145/356012.356014.
- [ChCo82] Chave, A. D., and C. S. Cox, 1982, Controlled electromagnetic sources for measuring electrical conductivity beneath the oceans: 1. forward problem and model study: Journal of Geophysical Research, 87, 5327–5338; DOI: 10.1029/JB087iB07p05327.
- [Ghos70] Ghosh, D. P., 1970, The application of linear filter theory to the direct interpretation of geoelectrical resistivity measurements: Ph.D. Thesis, TU Delft; UUID: 88a568bb-ebee-4d7b-92df-6639b42da2b2.
- [GuSi97] Guptasarma, D., and B. Singh, 1997, New digital linear filters for Hankel J0 and J1 transforms: Geophysical Prospecting, 45, 745–762; DOI: 10.1046/j.1365-2478.1997.500292.x.
- [HaJo88] Haines, G. V., and A. G. Jones, 1988, Logarithmic Fourier transformation: Geophysical Journal, 92, 171–178; DOI: 10.1111/j.1365-246X.1988.tb01131.x.
- [Hami00] Hamilton, A. J. S., 2000, Uncorrelated modes of the non-linear power spectrum: Monthly Notices of the Royal Astronomical Society, 312, pages 257–284; DOI: 10.1046/j.1365-8711.2000.03071.x; Website of FFTLog: casa.colorado.edu/~ajsh/FFTLog.
- [HuTS15] Hunziker, J., J. Thorbecke, and E. Slob, 2015, The electromagnetic response in a layered vertical transverse isotropic medium: A new look at an old problem: Geophysics, 80(1), F1–F18; DOI: 10.1190/geo2013-0411.1; Software: software.seg.org/2015/0001.
- [Key09] Key, K., 2009, 1D inversion of multicomponent, multifrequency marine CSEM data: Methodology and synthetic studies for resolving thin resistive layers: Geophysics, 74(2), F9–F20; DOI: 10.1190/1.3058434. Software: marineemlab.ucsd.edu/Projects/Occam/1DCSEM.
- [Key12] Key, K., 2012, Is the fast Hankel transform faster than quadrature?: Geophysics, 77(3), F21–F30; DOI: 10.1190/geo2011-0237.1; Software: software.seg.org/2012/0003.
- [Kong07] Kong, F. N., 2007, Hankel transform filters for dipole antenna radiation in a conductive medium: Geophysical Prospecting, 55, 83–89; DOI: 10.1111/j.1365-2478.2006.00585.x.
- [Shan55] Shanks, D., 1955, Non-linear transformations of divergent and slowly convergent sequences: Journal of Mathematics and Physics, 34, 1–42; DOI: 10.1002/sapm19553411.
- [SIHM10] Slob, E., J. Hunziker, and W. A. Mulder, 2010, Green's tensors for the diffusive electric field in a VTI half-space: PIER, 107, 1–20: DOI: 10.2528/PIER10052807.

- [Talm78] Talman, J. D., 1978, Numerical Fourier and Bessel transforms in logarithmic variables: Journal of Computational Physics, 29, pages 35–48; DOI: 10.1016/0021-9991(78)90107-9.
- [Tref00] Trefethen, L. N., 2000, Spectral methods in MATLAB: Society for Industrial and Applied Mathematics (SIAM), volume 10 of Software, Environments, and Tools, chapter 12, page 129; DOI: 10.1137/1.9780898719598.ch12.
- [Weni89] Weniger, E. J., 1989, Nonlinear sequence transformations for the acceleration of convergence and the summation of divergent series: Computer Physics Reports, 10, 189–371; arXiv: abs/math/0306302.
- [Wert17] Werthmüller, D., 2017, An open-source full 3D electromagnetic modeler for 1D VTI media in Python: empymod: Geophysics, 82(6), WB9–WB19; DOI: 10.1190/geo2016-0626.1.
- [Wert17b] Werthmüller, D., 2017, Getting started with controlled-source electromagnetic 1D modeling: The Leading Edge, 36, 352–355; DOI: 10.1190/tle36040352.1.
- [Wynn56] Wynn, P., 1956, On a device for computing the  $e_m(S_n)$  transformation: Math. Comput., 10, 91–96; DOI: 10.1090/S0025-5718-1956-0084056-6.
- [ZiSl19] Ziolkowski, A., and E. Slob, 2019, Introduction to Controlled-Source Electromagnetic Methods: Cambridge University Press; ISBN: 9781107058620.

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